

CHAOS in SIBERIA

A story for history

Short (40') account of a long (40 yrs) work

Boris Chirikov & many coworkers in

Budker Institute of Nuclear Physics

& elsewhere

My (accidental) start

(≈ 1958, Kurchatov Institute, Moscow)

a simple-looking but turned-out rich

Budker's problem: single particle confinement in

Budker's (adiabatic) magnetic trap

(for the great END – controlled nuclear fusion !)

Toulouse

16 July 1998

BACKGROUND (retrospectively)

Revival of intensive studies
into nonlinear dynamics,
surprising rediscovery of chaos (stochasticity) after
Boltzmann ... Poincare ...

- new applications:
strong focusing accelerators,
controlled nuclear fusion
- computers, NUMERICAL EXPERIMENTS !!

– Lehmer, 1951

pseudorandom number generators after
Galton Board (= Lorentz gas in external field !)

– Goward and Hine (CERN), 1953, accelerators

➔ Fermi, Pasta and Ulam, 1955

foundations of statistical mechanics

– Symon and Sessler, 1956, accelerators

➔ – Kolmogorov, 1954, KAM-integrability
(in spite of Poincare theorem !)

Budker's problem:

2-freedom nonlinear oscillations
with a weak adiabatic coupling

$$(e = m = c = 1)$$

$$W = \frac{v_{\parallel}^2}{2} + \frac{v_{\perp}^2}{2} = \frac{p^2}{2} + \mu \cdot \omega(s) = \text{const}$$

$$\omega(s) = \omega_0 \left(1 + \frac{s^2}{2L^2} + \dots \right) = \langle \omega \rangle + \tilde{\omega}(t)$$

frequency modulated Larmor spiralling

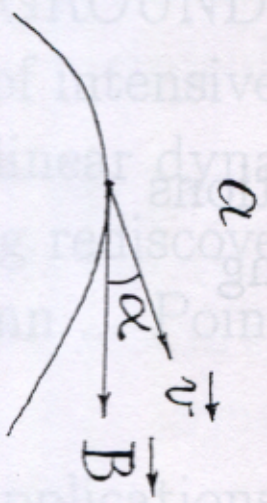
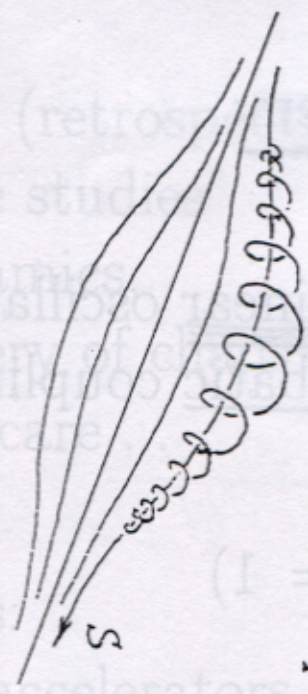
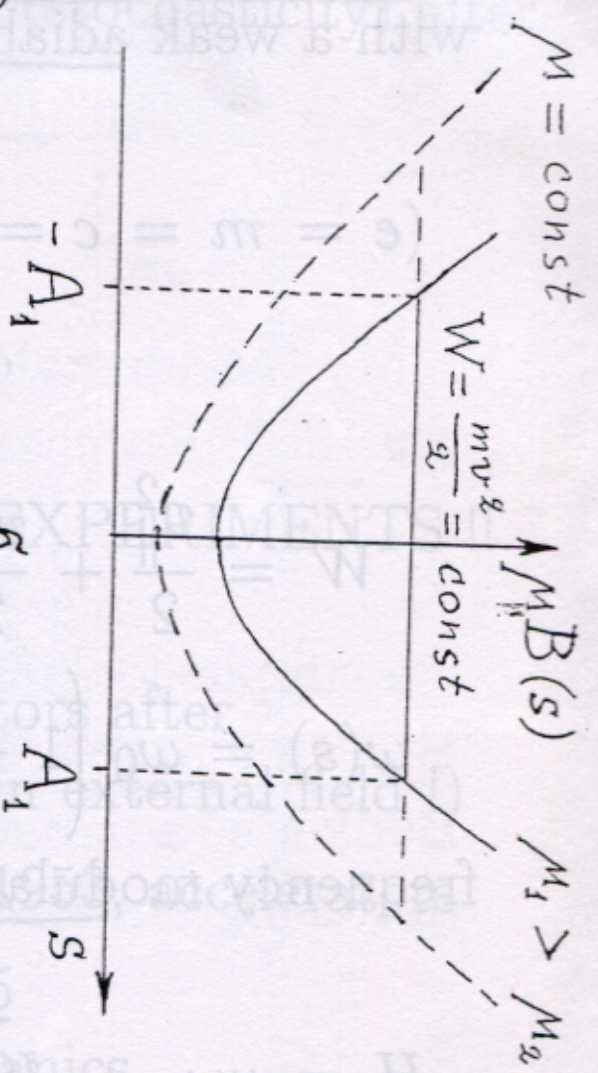
$$H = \mu\omega_0 + J\Omega(\mu) + V_a$$

adiabatic

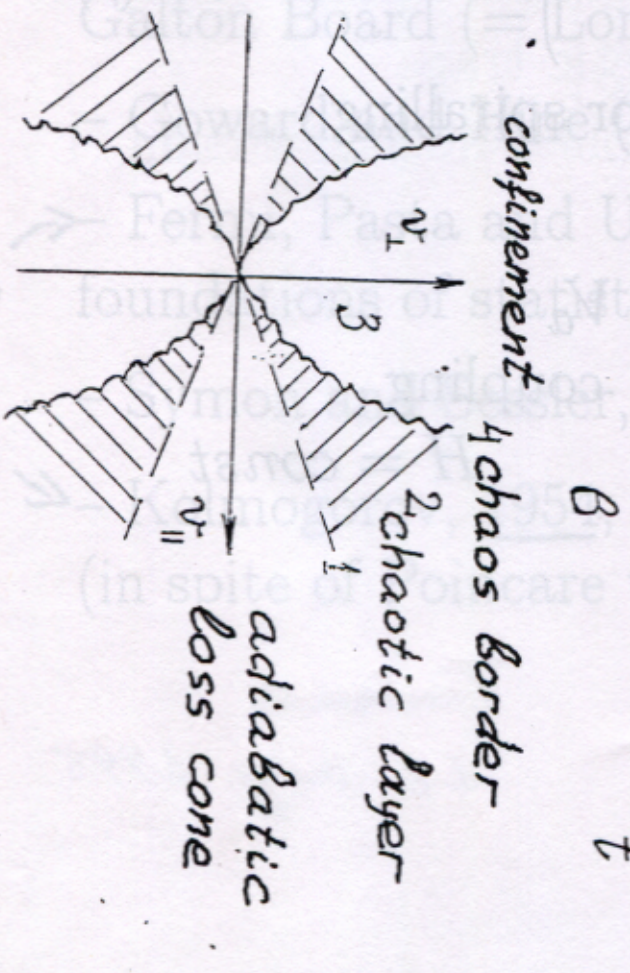
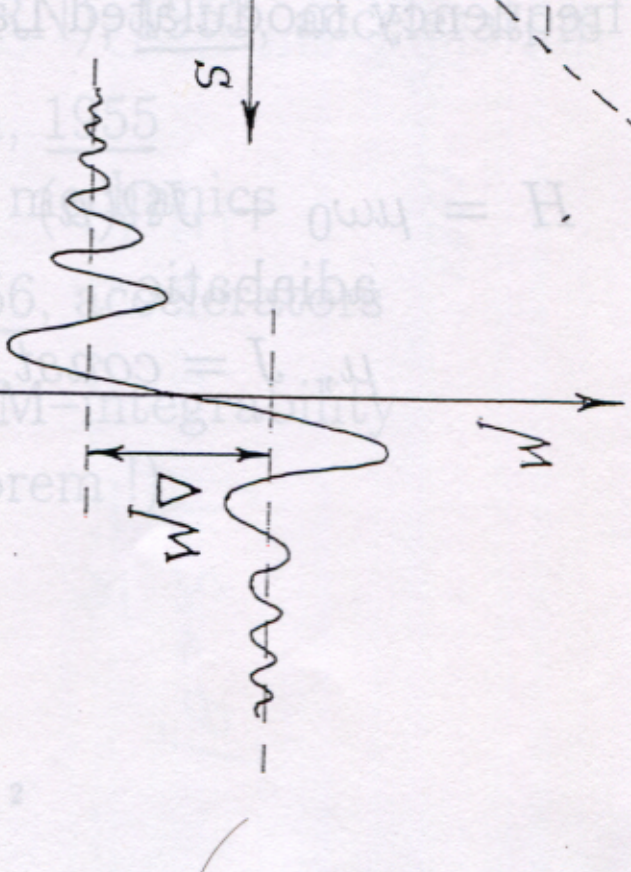
coupling

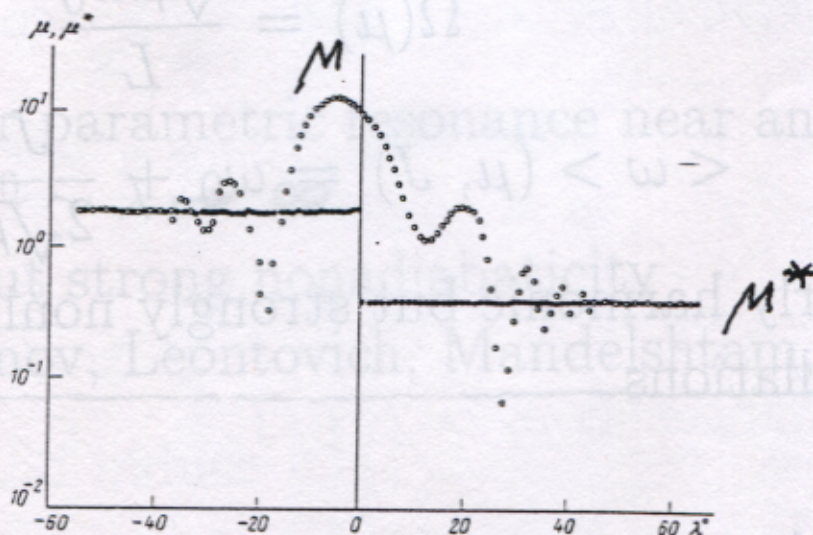
$$\mu, J = \text{const}$$

$$H = \text{const}$$



$$M \doteq \frac{mv^2}{2B(s)} \cdot \sin^2 \alpha(s)$$





Зависимость $\mu \sim \sin^2 \alpha / B$ (белые кружки) и $\mu^* \sim \sin^2 \alpha^* / B$ (черные кружки) от широты λ при движении частицы с энергией $W = 200$ МэВ между точками отражения в геомагнитной дипольной ловушке. На экваторе ($\lambda = 0$): R_{eff} (в радиусах Земли) $\approx 2,9$ ($\chi_i = 0,272$); $\alpha = 20^\circ$; $\phi = 108^\circ$

$$\alpha_n^*, \phi_n \rightarrow \alpha_{n+1}^*, \phi_{n+1}, \quad (5)$$

где $\alpha_{n+1}^* = \alpha_n^* + \Delta \alpha_n^* \cos \phi_n$, $\phi_{n+1} = \phi_n + \Delta \phi(\alpha_{n+1}^*)$. Изменение фазы (набег фазы) между двумя последовательными пересечениями экваториальной плоскости приближенно описывается выражением

$$\Delta \phi = \frac{\pi \bar{\omega}}{\Omega} = \frac{6F(\alpha)}{\chi}, \quad F(\alpha) \approx \sin^{-1,348} \alpha - 0,255, \quad 1,25^\circ \leq \alpha \leq 90^\circ, \quad (6)$$

где $\bar{\omega}$ — ларморовская частота, усредненная по продольному колебанию, Ω — частота продольных осцилляций. Окончательно с учетом (3)–(6) приходим к отображению

$$\cos \alpha_{n+1}^* = \cos \alpha_n^* \cos \nu + \sin \alpha_n^* \sin \nu \cos(\phi_n - \bar{\phi}_0), \quad (7)$$

$$\phi_{n+1} = \phi_n + \frac{6F(\alpha_{n+1})}{\chi_i}$$

где $\bar{\phi}_0 = \arcsin(\sin \alpha_0 \sin 2\delta / \sin \nu)$. По аналогии с ¹ рекуррентному соотношению (5) можно придать вид стандартного отображения Чирикова

$$I_{n+1}^* = I_n^* + K^* \sin \theta_n, \quad \theta_{n+1} = \theta_n + I_{n+1}^*, \quad (8)$$

где $I^* = \frac{\partial}{\partial \alpha} \left(\frac{\partial F}{\partial \alpha^*} \right)_{\alpha^*} \cdot (\alpha^* - \alpha_r^*)$, $K^* = \frac{6\nu}{\chi} \left(\frac{\partial F}{\partial \alpha^*} \right)_{\alpha^*}$, $\theta = \phi - \bar{\phi}_0 - \pi/2$, α_r^* находится из условия резонанса $\bar{\omega} = 2r\Omega$ (r — любое целое число). В итоге получается, что рассматриваемая модель движения включает в себя "адиабатическое" движение ($\mu^* = \text{const}$) между экватором и точкой отражения и дискретную модель (7) (или 8) для многократных продольных колебаний частицы. Данная

- nonlinearity

$$\Omega(\mu) = \frac{\sqrt{\mu\omega_0}}{L}$$

$$\langle \omega \rangle (\mu, J) = \omega_0 + \frac{J}{2\sqrt{\mu}} \cdot \frac{\sqrt{\omega_0}}{L}$$

nearly harmonic but strongly nonlinear oscillations

- adiabaticity

(big) adiabaticity parameter

$$\lambda \sim \frac{\langle \omega \rangle}{\Omega} \sim \frac{L}{\rho} \gg 1$$

a single swing:

$$\Delta\mu \sim e^{-\lambda} \rightarrow \exp\left(-\frac{2\sqrt{2}}{3} \cdot \frac{L\omega_0}{v}\right)$$

(nonperturbative theory)

• RESONANCES

a linear parametric resonance near any integer

$$r = \frac{2\omega_0}{\Omega} \sim \lambda \rightarrow \infty$$

slow but strong nonadiabaticity

Andronov, Leontovich, Mandelshtam, 1928

• NONLINEAR RESONANCES ???

Poincare map (standard) $(p, x) \rightarrow (\bar{p}, \bar{x})$:

$$\bar{p} = p + k \cdot \sin x$$

$$\bar{x} = x + T \cdot \bar{p}$$

with single parameter $K = kT$

$$H(x, p, t) = T \frac{p^2}{2} + k \cdot \sum_{r=-\infty}^{\infty} \cos(x - 2\pi r t)$$

adiabatic

perturbation

$p = \text{const}$

resonances at $p_r = 2\pi r/T$

r integer

a single nonlinear resonance ('pendulum')

width $\Delta p = 4\sqrt{k/T}$, $\Delta\omega = 4\sqrt{K}$, ($\omega \approx \dot{x}$)
 (a new) frequency $\Omega_r = \sqrt{K} \ll 1$

completely integrable

suppresses resonant perturbation

(a big hope failed !)

INTERACTION of RESONANCES

2 (or more) resonances

with spacing $\delta p = 2\pi/T$, $\Delta\omega = 2\pi$

interact and induce

a NEW INSTABILITY

surprisingly irregular, chaotic ?

a natural (overlap) parameter

$$\boxed{S \sim \frac{\Delta p}{\delta p}} = \frac{\Delta\omega}{\delta\omega} = \frac{2}{\pi} \sqrt{K} \sim 1$$

eternal particle confinement for some $K \rightarrow 0$

Arnold, 1962

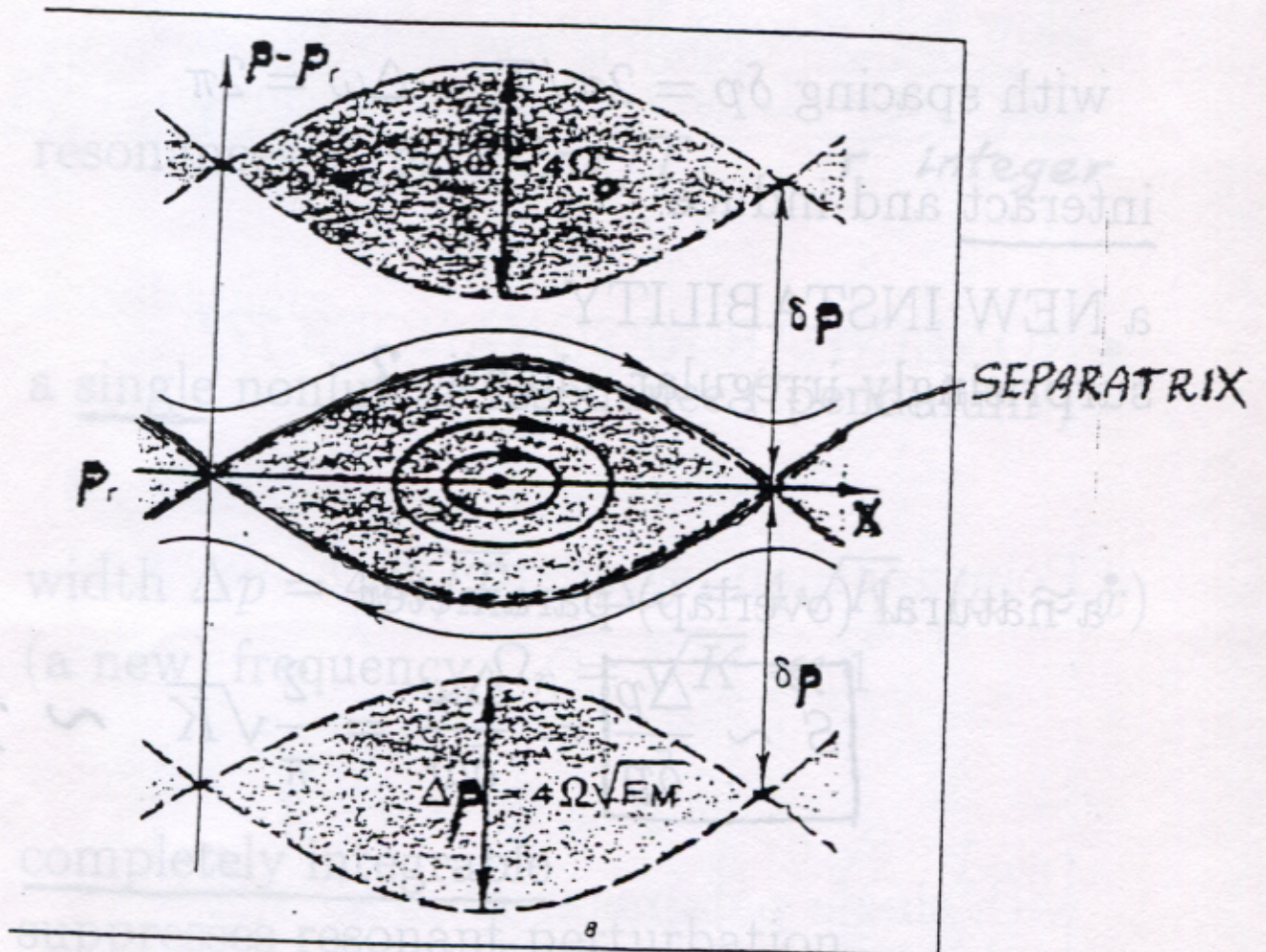
3F (asymmetric) trap ?

the universal instability, Arnold, 1964

Arnold diffusion

$$H(x, p, t) = T \frac{p^2}{2} + k \cdot \sum_{r=-\infty}^{\infty} \cos(x - 2\pi r t)$$

Б. В. Чириков



Значение величин $\dot{\theta}$ и θ в начальный момент времени, для которых имеют место

$S \gg 1$ apparently global instability,
surprisingly irregular, chaotic?

KS-entropy

$$h = \Lambda > 0$$

Kolmogorov, 1958

Sinai, 1962

diffusion

$$D = \frac{\langle (\Delta p)^2 \rangle_t}{t} = \frac{k^2}{2} C(K) \xrightarrow{K \rightarrow \infty} \frac{k^2}{2}$$

Rechester, Rosenbluth, White, 1981

no confinement (?)

$S \ll 1$ nearly (but not completely!) integrable,
apparently stable,
confinement (?)

②F (axisymmetric trap):

eternal particle confinement for some $K \rightarrow 0$

Arnold, 1962

③F (asymmetric) trap ?

the universal instability, Arnold, 1964

Arnold diffusion

SEPARATRIX: most unstable place,
the universal (for any $K \rightarrow 0$) origin of chaos,
chaotic layer *etc*

CROSS-ROAD  other sophisticated applications
to the great ENDS
simple models for
fundamental laws (of the random)

STRATEGY (spontaneous)

a physical RESONANCE theory

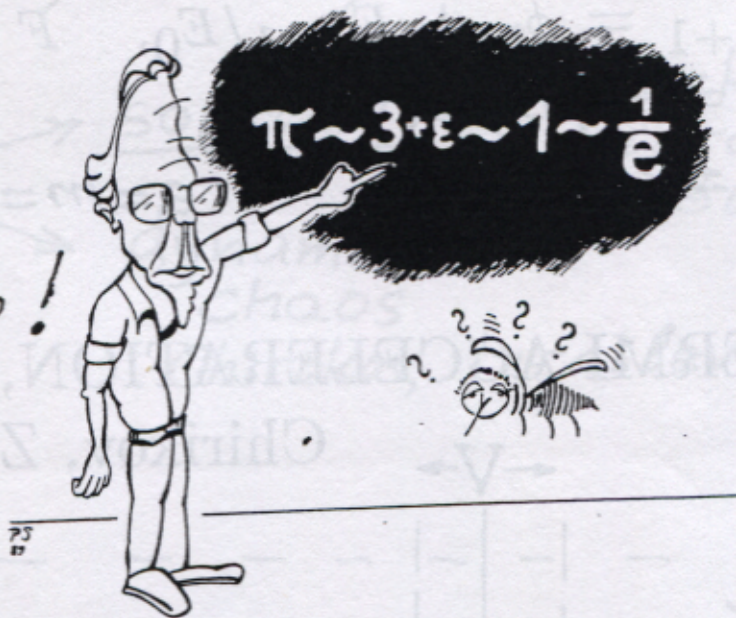


NUMERICAL EXPERIMENTS !!

A physical theory...

$$\pi \sim 3 + \varepsilon \sim 1 \sim \frac{1}{e}$$

Chercher
la resonance!



COMET HALLEY

diffusion backwards in time

$t \sim 10 \text{ Myr}$

Chirikov, Vecheslavov, 1989

instructive
examples

SIMPLE CHAOS:

most complex dynamics

BUT simple statistics

- MICROTRON, Veksler, 1944

$$E_{n+1} = E_n + V_0 \sin \phi_n$$

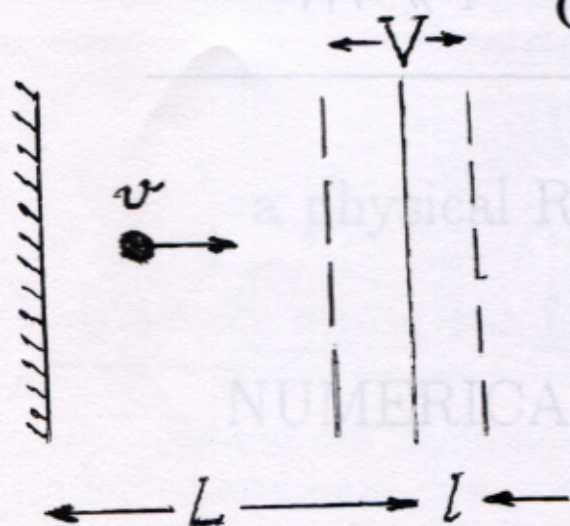
$$\phi_{n+1} = \phi_n + E_{n+1}/E_0, \quad F = 2 \text{ freedoms}$$

$$K = V_0/E_0 \gg 1$$

$$E_0 = \omega_0/2\pi\Omega, \quad e = m = c = 1$$

- FERMI ACCELERATION, Ulam, 1961

Chirikov, Zaslavsky, 1964



$$\frac{v^2}{V^2} < \left(\frac{4L}{l} \right) !$$

$$(\text{adiabatics: } \frac{v^2}{V^2} \lesssim \left(\frac{L}{l} \right)^2)$$

- COMET HALLEY

diffusion backwards in time

$$t_L \sim 10 \text{ Myrs} \quad ?$$

Chirikov, Vecheslavov, 1989

Fermi's problem (1923)

ergodicity of nonlinear oscillation:

$$H = \sum_{i=1}^N \frac{\dot{x}_i^2}{2} + \frac{(x_i - x_{i-1})^2}{2} + (x_i - x_{i-1})^p, \quad p=3;4, \quad N \gg 1$$

Fermi, Pasta, Ulam, 1955

nonergodic (generally) $\rightarrow$ solitons
 $\rightarrow$ dynamical chaos

Kruskal,
Miura,
Zabusky, 1966

Chirikov, Izrailev, 1966

De Luca, Lichtenberg, Lieberman, 1995
Shepelyansky, 1997

Fermi, Pasta, Ulam, 1955

NO RELAXATION ???

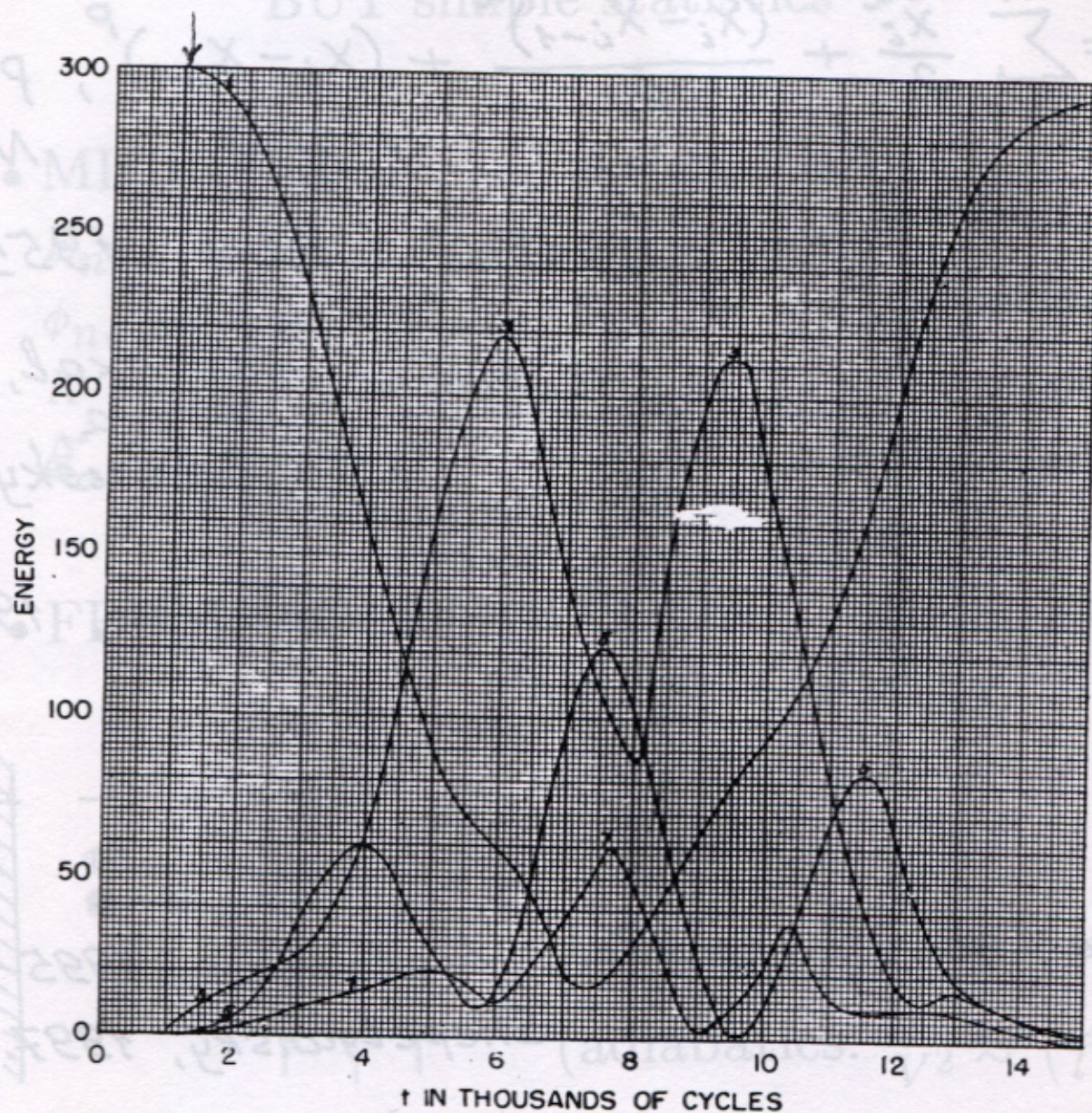


Fig. 4. The initial configuration assumed was a single sine wave; the force had a cubic term with $\beta = 8$ and $\delta t^2 = 1/8$. Since a cubic force acts symmetrically (in contrast to a quadratic force), the string will forever keep its symmetry and the effective number of particles for the computation $N = 16$. The even modes will have energy 0.

Fermi, Pasta, Ulam, 1955
RELAXATION !!!

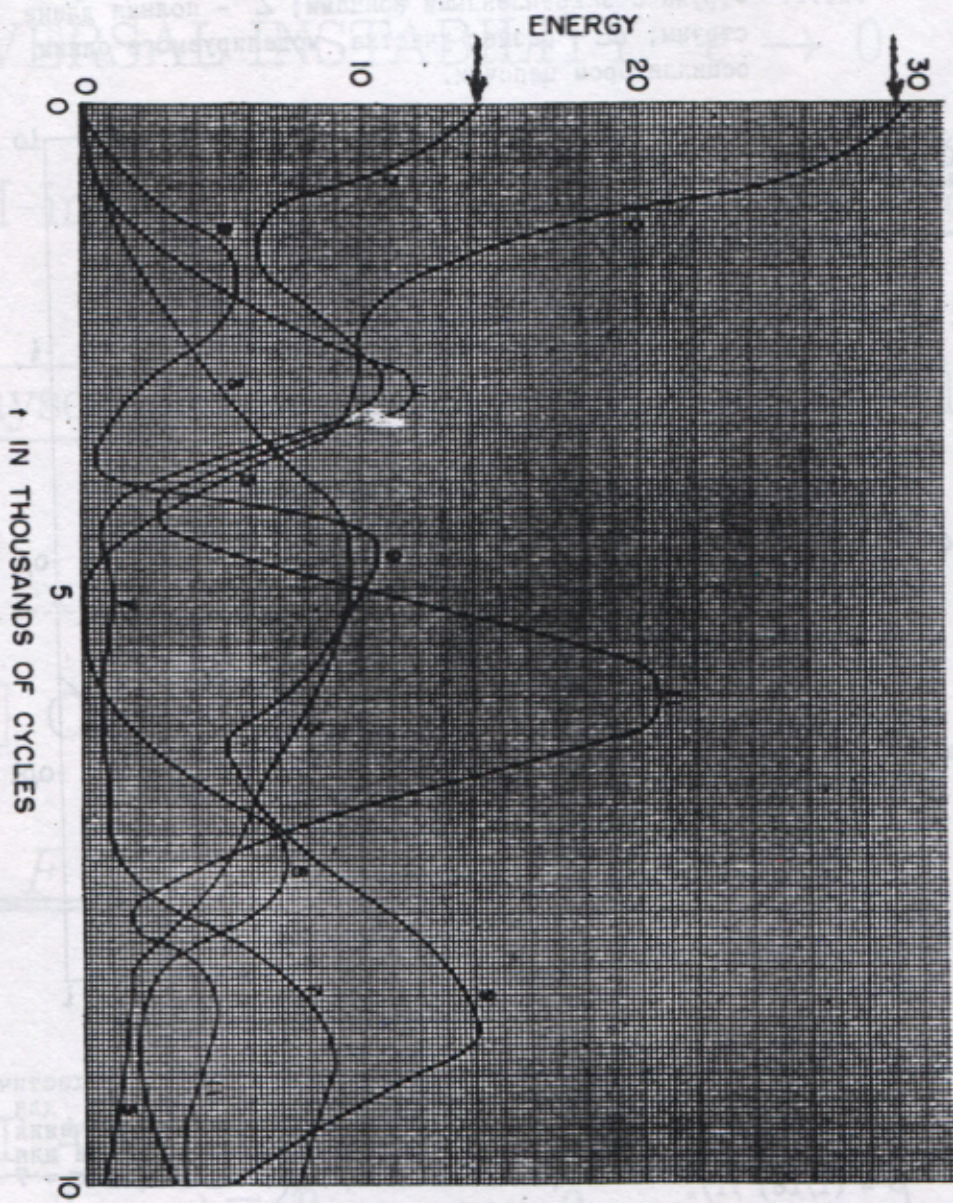


Fig. 5. ($N = 32$) $\delta t^2 = 1/64$; $\beta = 1/16$. The initial configuration was a combination of 2 modes. The initial energy was chosen to be $2/3$ in mode 5 and $1/3$ in mode 7.

Fermi, Pasta, Ulam, 1955

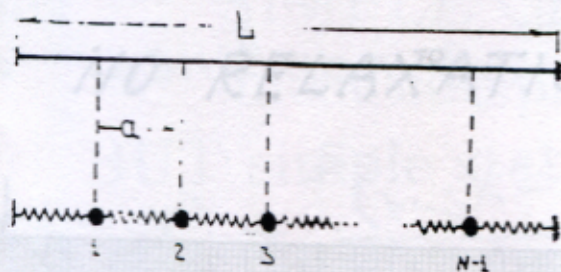


Рис.1. Струна с закрепленными концами; L - полная длина струны; a - размер участка, моделируемого одним осциллятором цепочки.

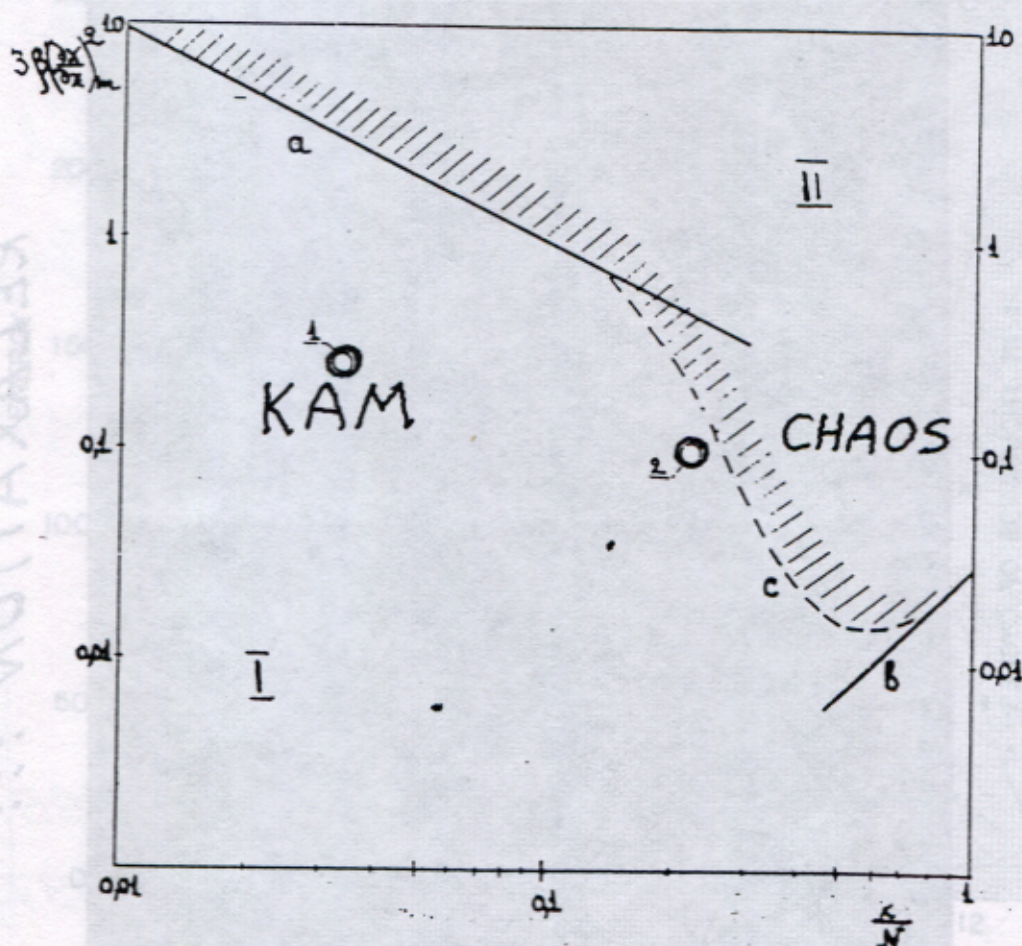


Рис.2. I - область колмогоровской устойчивости; II - область стохастичности
 a - граница стохастичности для $k \ll N$ (2.II) b - граница для $k \approx \sqrt{N}$ (3.5); c - качественная интерполяция; численные значения прямых a , b даны для $N = 32$; 1 - результат численного счёта для $N = 32$; $x_0 = 1$; $K = 1$; $\beta = 8/11$; 2 - тоже самое для $K = 7$; $\beta = (1/16) L/1$.

Fig. 4. The initial configuration assumed was a single sine wave; the force had a cubic term with $\beta = 8$ and $\epsilon/2 = 1/8$. Since a cubic force acts symmetrically (in contrast to a quadratic force), the string will forever keep its symmetry and the effective number of particles for the computation $N = 16$. The even nodes will have energy 0.

COMPLEX CHAOS:

most complex dynamics AND statistics

● ARNOLD DIFFUSION, 1964

$F > 2$ everywhere dense web of measure

$$\mu \sim \exp(-A \epsilon^{-1/2}), \quad D \sim \mu^2, \quad \lambda \sim \epsilon^{-1/2}$$

UNIVERSAL INSTABILITY $\epsilon \rightarrow 0$

and

KAM-integrability (= adiabatic invariance)

- Chirikov, Gadiyak, Izrailev, 1975
- [Tennyson], Lichtenberg and Lieberman, 1979
- Nekhoroshev, 1977
- $\mu \sim \exp(-A \epsilon^{-1/2L}), \quad 1 < L(F) = ?$
- [Ford], Chirikov and Vivaldi, 1979
- Lochak, 1990
- $L = F - 1$ for a simple resonance
in conservative systems

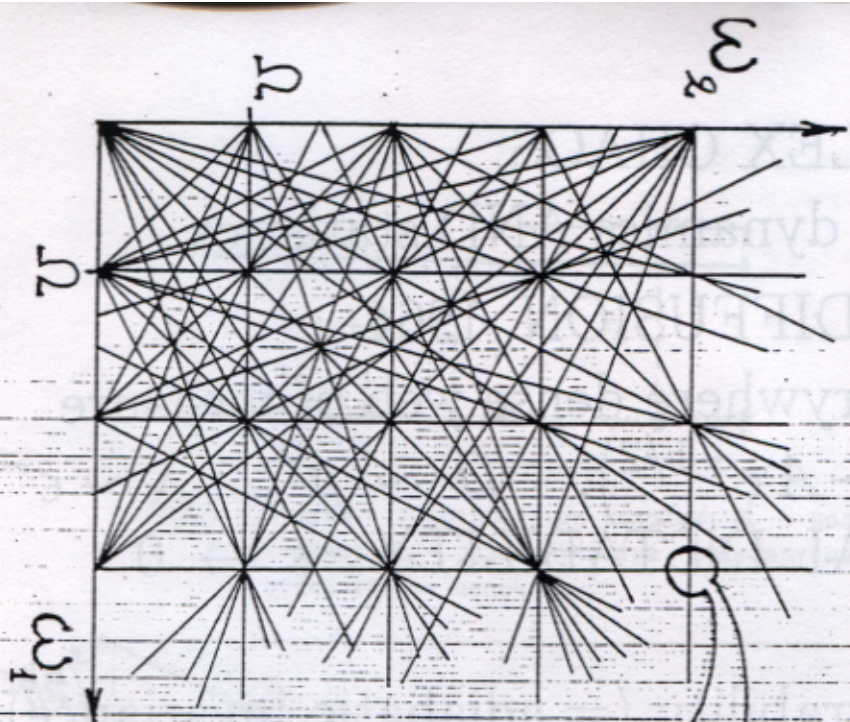
FAST ARNOLD DIFFUSION

$$\mu \sim \epsilon^p \sim \lambda^{-2p}, \quad p \approx 3$$

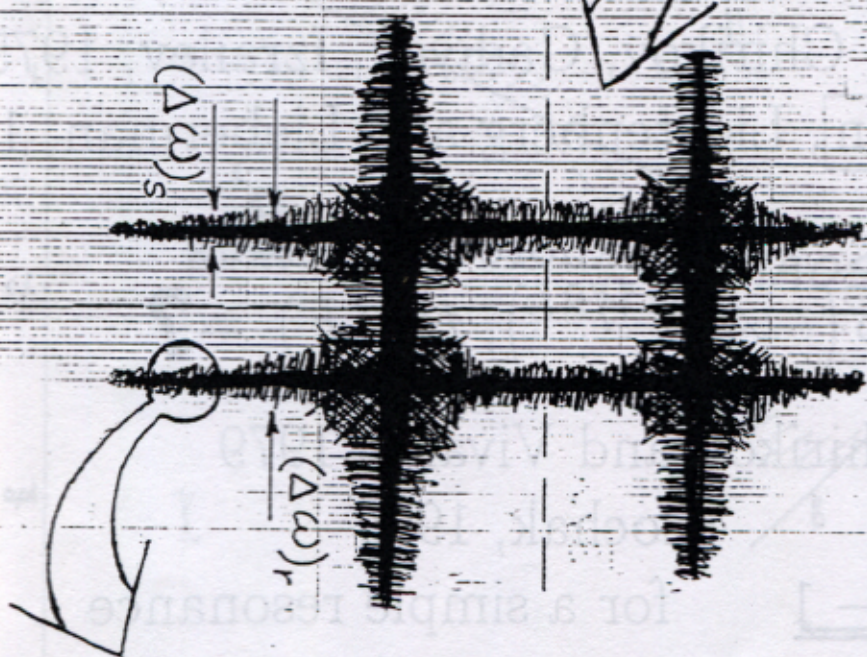
in intermediate asymptotics $1 \gg \epsilon \gtrsim e^{-2L}$

- Chirikov, Vecheslavov, 1996

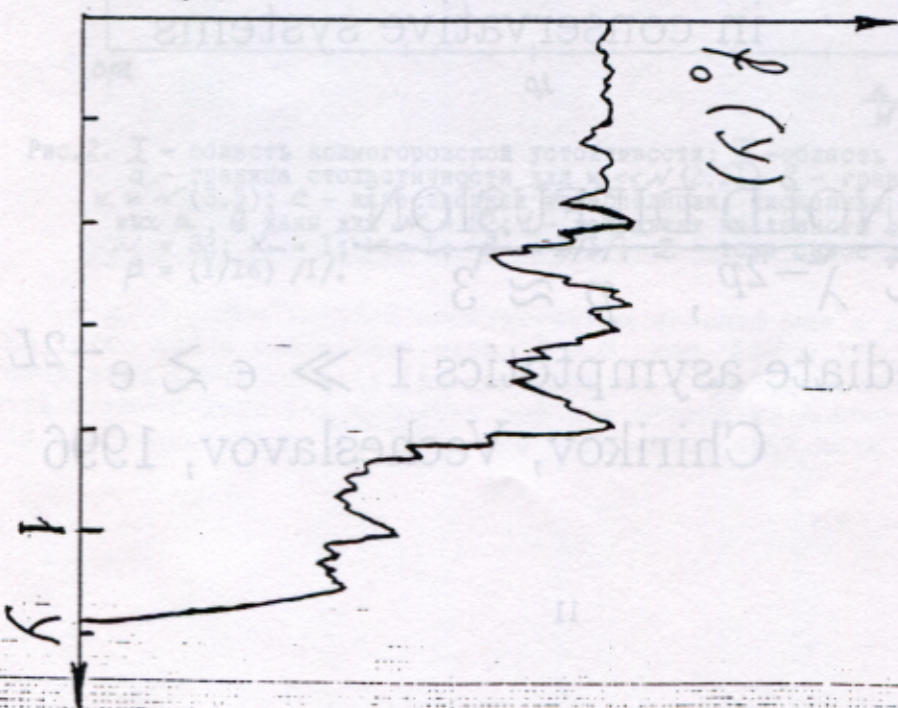
a



b



c



Kaneko, Konishi, 1993

Arnold web

$K = 0.2$, $N = 4$, $F = 3$

0.5

P2

-0.5

-0.5

P1

0.5

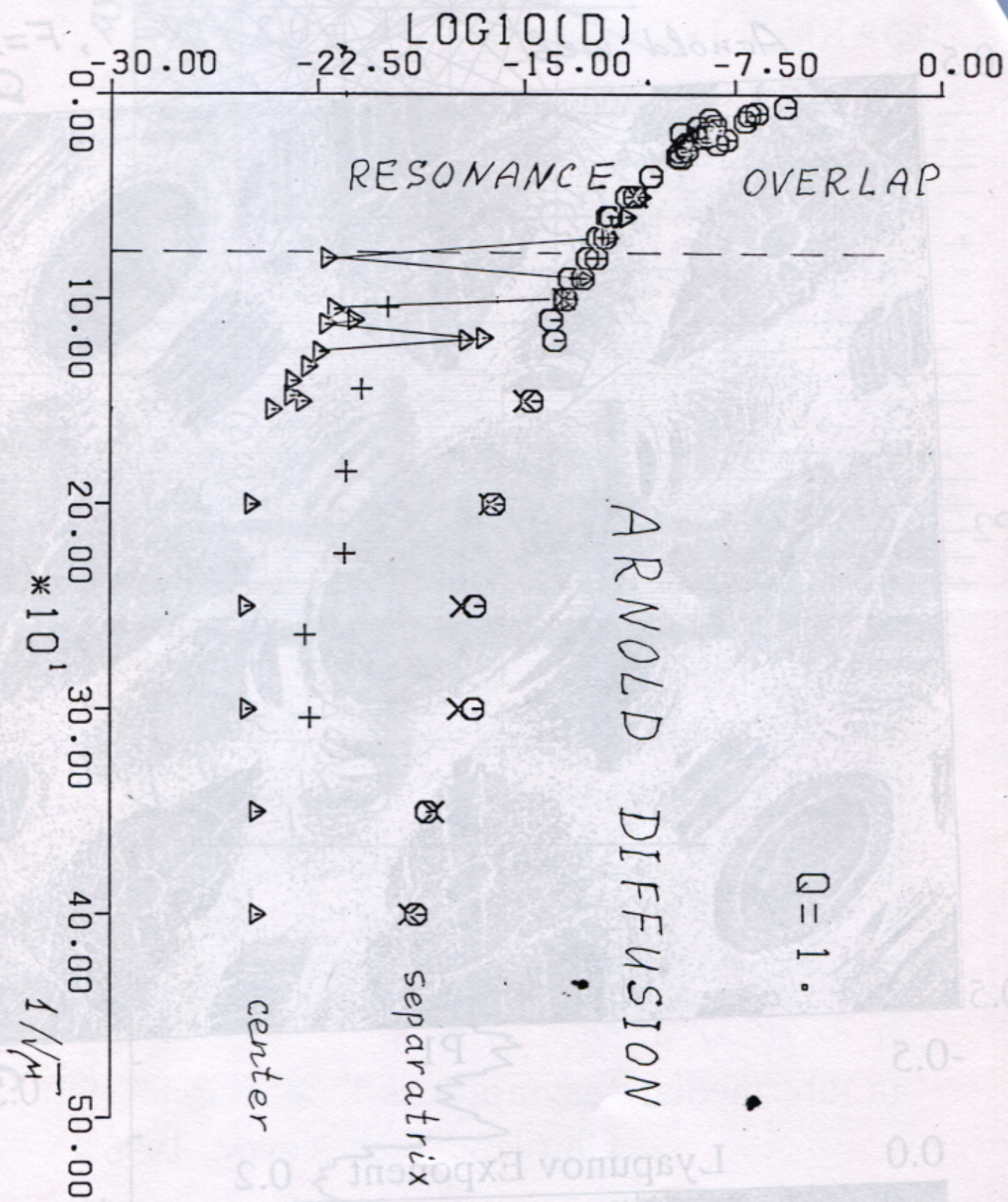
0.0

Lyapunov Exponent

0.2

$$H = \frac{|P|^2}{2} - K \sum_{i=1}^N \cos(x_{i+1} - x_i) \cdot \delta_1(t)$$

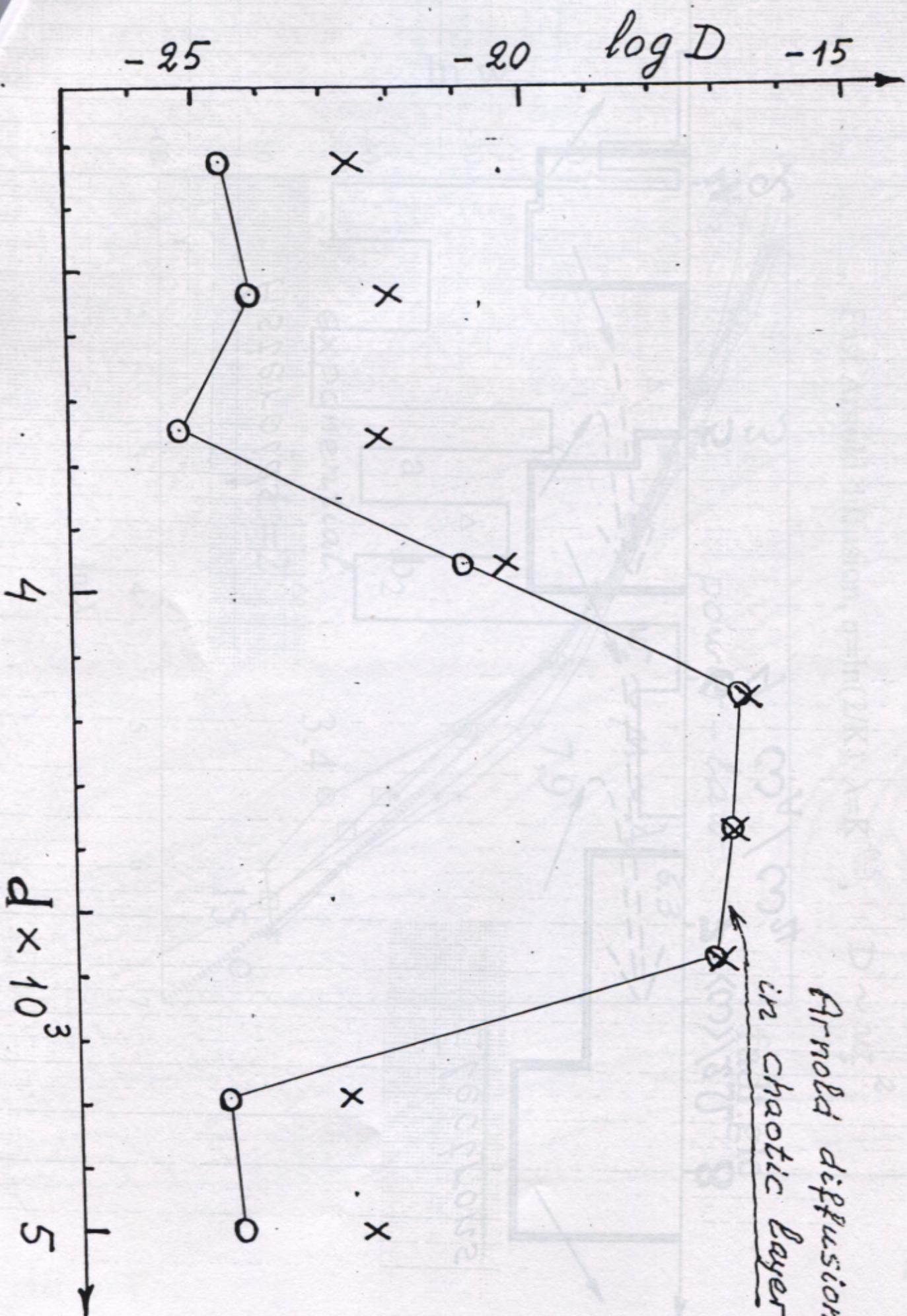
Chirikov, Ford, Vivaldi, 1973



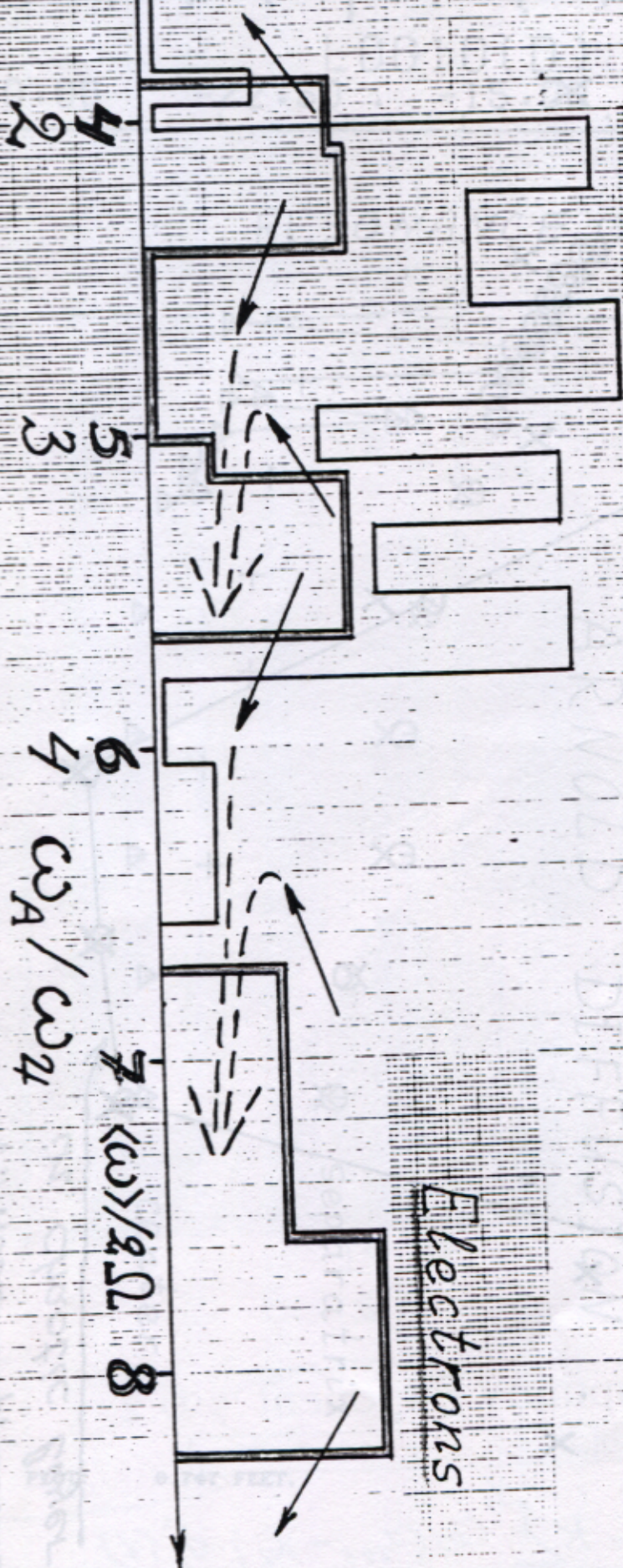
PLOT: 0.747 FEET. 100.000% SCALE.

Chirikov, Ford, Vivaldi, 1979

Arnold diffusion
in chaotic layer



Asteroids

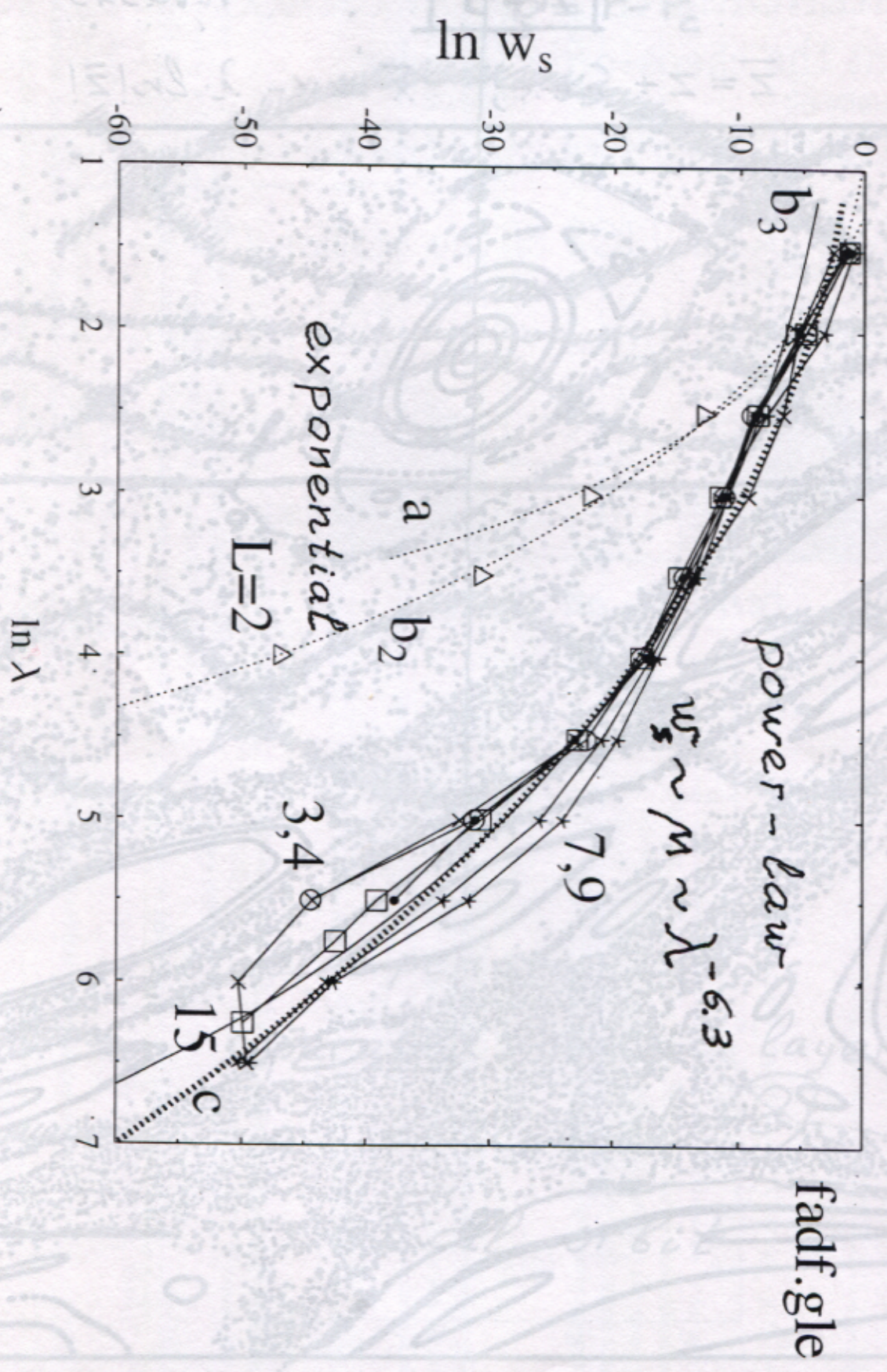


Electrons

$$\omega_A / \omega_Z$$

$$\langle \omega \rangle / 2\Omega$$

Fast Arnold diffusion, $\sigma = \ln(2/K)$, $\lambda = K^{-0.5}$, $D \sim w_s^2$



$$\boxed{\lambda = 5}$$

Vivaldi

$$\bar{z} = z + \sin x; \quad \bar{x} = x - \lambda \ln |\bar{z}|$$

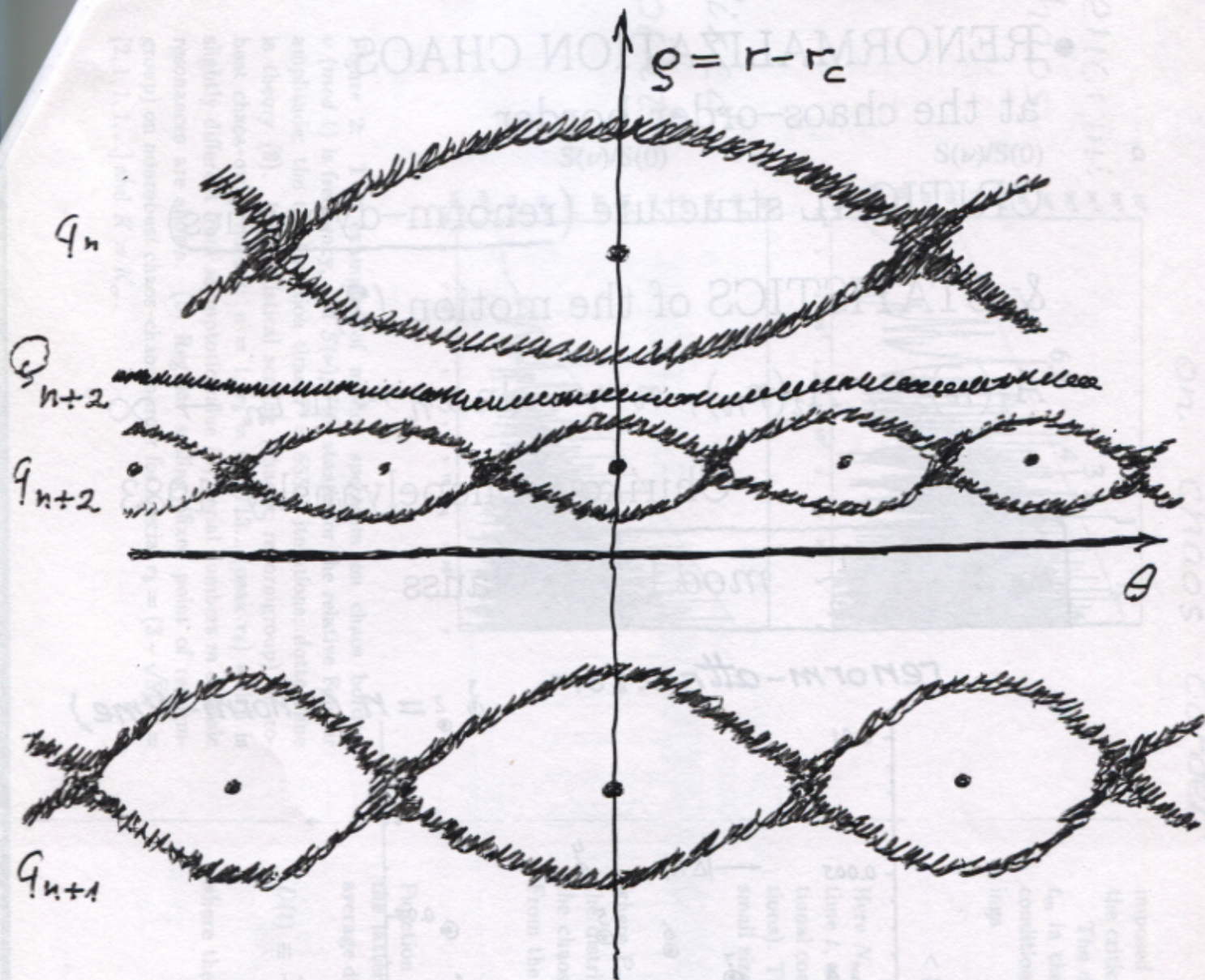


main chaos border — X

Fig 3b

Critical

structure



• stable periodic orbits

separatrix chaotic layers

②

minima

a bottleneck ②

critical orbit

- RENORMALIZATION CHAOS
at the chaos-order border

CRITICAL structure (renorm-dynamics)

& STATISTICS of the motion (?)

$$A_i(n) \rightarrow A_i(r_n), \quad n \sim -\ln x_n \sim \ln t_n \rightarrow \infty$$

Chirikov, Shepelyansky, 1983

$$r_{n+1} = \frac{1}{r_n} \bmod 1 \quad \text{Gauss}$$

renorm-attractor

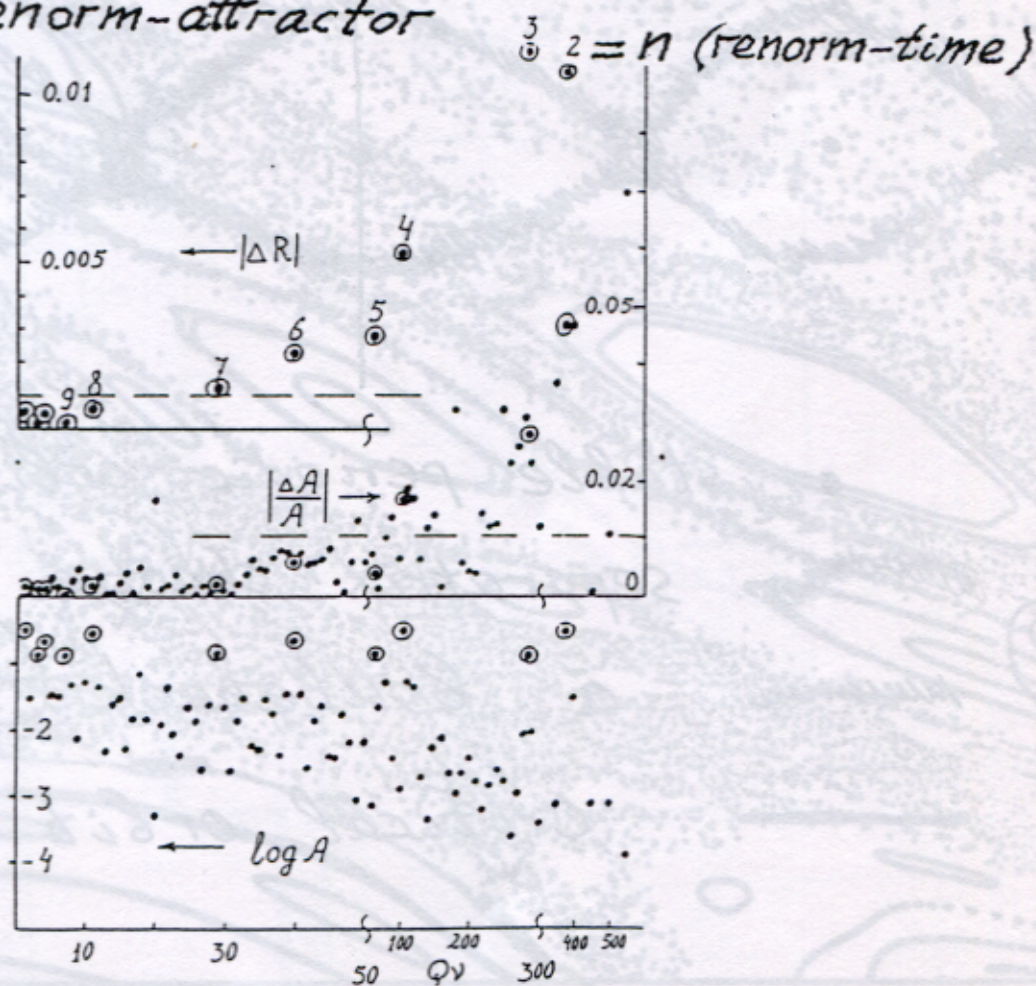


Fig. 4. Chaotic "renormalization attractor" in standard map (see the text). Arrows show the related scales. Notice changes in scale along QV axis. Numbers at upper points are the values of renormalization time n . The principal scales are marked by circles.

Fig 3b

Motion spectrum on chaos border

renormal-
chaos

fixed
point

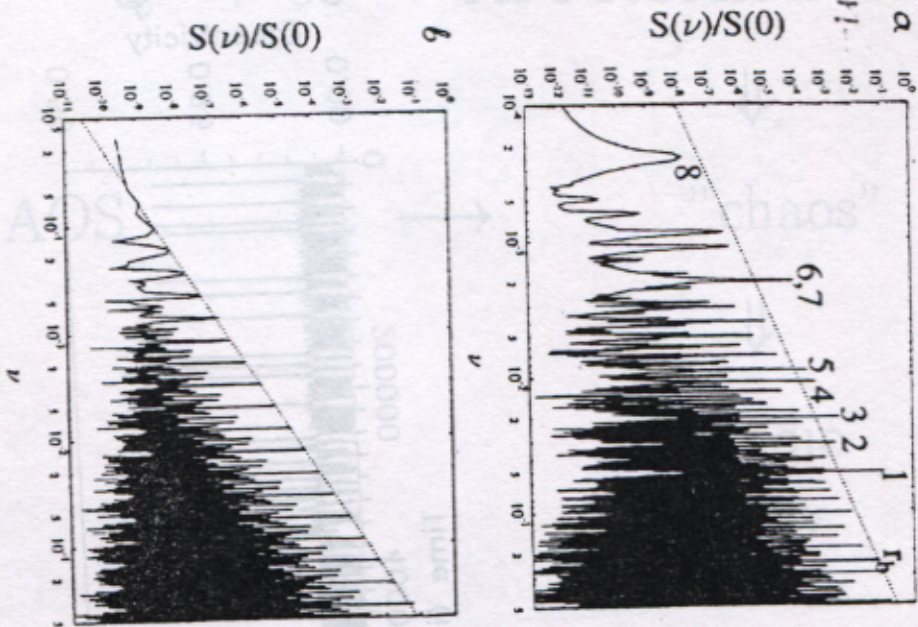


Figure 2: Two examples of motion spectrum on chaos border: ν (*mod* 1) is frequency, and $S(\nu)/S(0)$ stands for the relative Fourier amplitude; the total motion time $T = 65536$ iterations; dotted line is theory (8). (a) Statistical scaling (chaotic renormgroup) on robust chaos-order border: $n = 1$; $r_b = 0.23713\dots$ (peak r_b) which is slightly different from asymptotic value (6); serial numbers m of basic resonances are shown. (b) Regular scaling (fixed point of renorm-group) on nonrobust chaos-chaos border for special $r_b = (3 - \sqrt{5})/2 = [2, 1, 1, 1, \dots]$ and $K = K^{cr}$.

imposed on strong fluctuations which are generally characteristic for the critical phenomena.

The diffusion rate is determined by the statistic of sticking times t_m in the corresponding scale m . In average over time or the initial conditions (ergodicity), and assuming statistical independence of stickings

$$\langle (\Delta p)^2 \rangle \approx \sum_m (\Delta p)_m^2 \approx K^2 \sum_m t_m^2 N_m + \frac{K^2}{2} C_0(K) t \quad (9)$$

Here $N_m(t)$ is the number of entries into scale m for the total motion time t , and the latter term describes the normal diffusion (with additional coefficient $C_0(K) \sim 1$ which accounts for the short-time correlations). The normal diffusion occupies the most of the time owing to a small size of regular domains. In turn, the number of entries

$$N_m = t \cdot P_m, \quad P_m \sim \frac{A}{t_m^c} \quad (10)$$

where $P_m = P(t_m)$ stands for the Poincare recurrence statistic that is the distribution of the delays during the reflection (scattering) from the chaos border which is characterized by the critical exponent c_P . From the motion ergodicity

$$\frac{t_m N_m}{t} = t_m P_m = A_m \sim \frac{A}{q_m^2} \sim \frac{A}{t_m^{c_A}} \quad (11)$$

Function $A_m = A(t_m)$ plays a role of the sticking correlation, and from the latter estimate (scaling) $c_P = c_A + 1$. Whence, the asymptotic average diffusion rate

$$D(t) \equiv \frac{\langle (\Delta p)^2 \rangle}{t} \rightarrow K^2 \sum_m t_m A_m \sim AK^2 t_{max}^{1-c_A} \sim AK^2 \cdot (At)^{c_D} \quad (12)$$

where the critical diffusion exponent

$$c_D = \frac{1 - c_A}{1 + c_A} \quad (13)$$

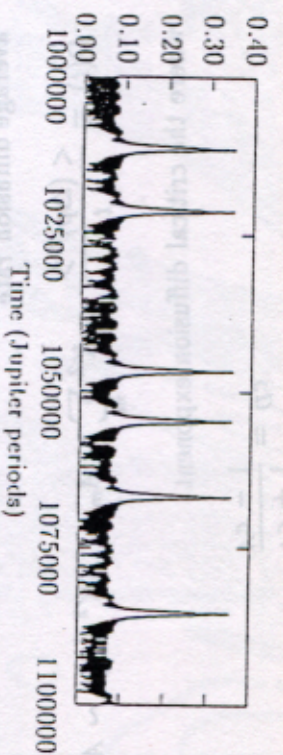
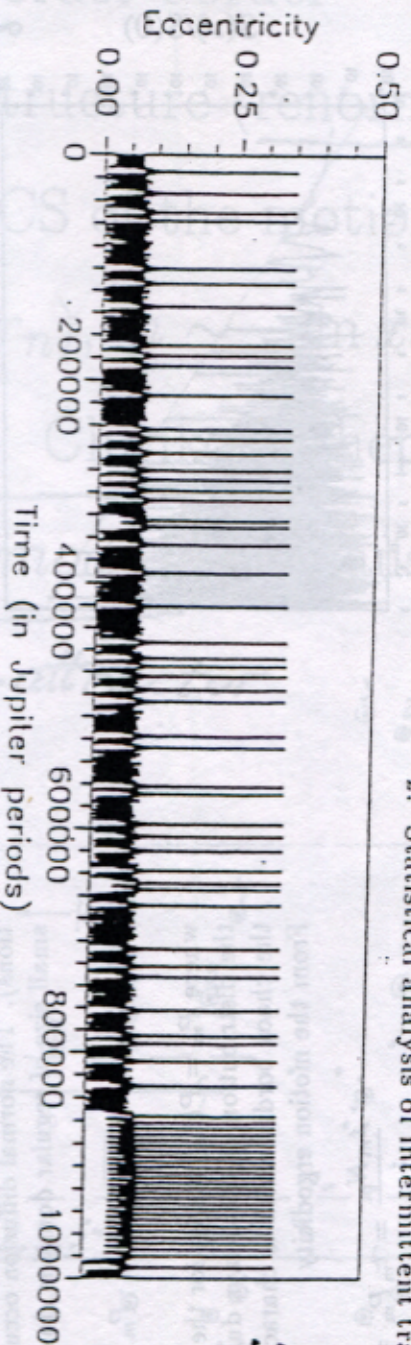


Figure 1. Eccentricity versus time for an intermittent trajectory.

2. Statistical analysis of intermittent trajectories

*asteroid in
3/1 resonance*



QUANTUM $\longrightarrow$ classical
 mechanics $Q \rightarrow \infty$ APPROXIMATION

$\Downarrow$ $\Downarrow$
 PSEUDOCHAOS $\longrightarrow$ "chaos"
 $\Downarrow$ $\Downarrow$
 ? ergodic theory

COMPUTER PSEUDOCHAOS

Budker's problem: single particle confinement in
 Budker's (adiabatic) magnetic trap
 (for the great END - controlled nuclear fusion !)

1824
 15 10
 11 11
 5 0
 + +
 [k·2·5]
 15
 [0]

Toulence
 16 July 1998

B. V. Chirikov, Izrailev, Shepelyansky, 1981

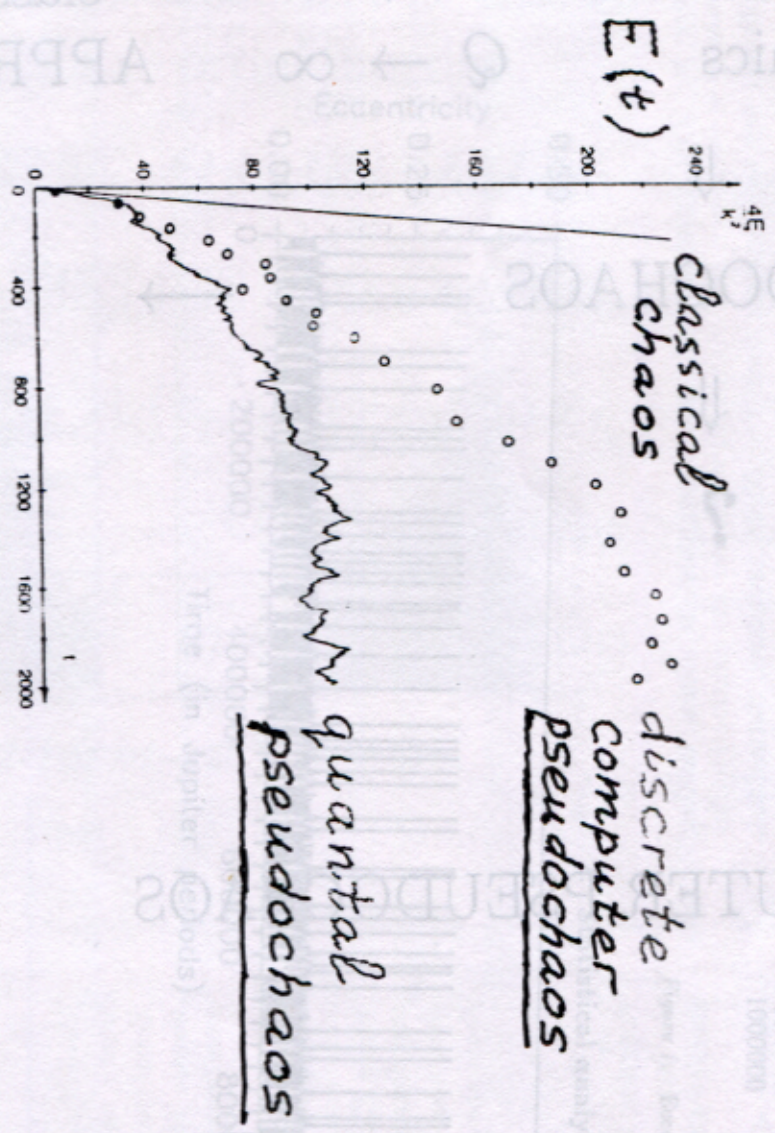


Fig. 21. "Quantum localization" in the classical discrete map (7.13) [38]: the straight line represents classical diffusion, the wiggly curve shows quantum localization, and circles are for the discrete model; $k = 20$; $N = 5$.

$$\bar{n} = n + [k \cdot \sin \theta]$$

$$\bar{\theta} = \theta + T \cdot \bar{n}.$$

3/1 resonance