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Dynamics and thermalization of Bose-Einstein condensate in Sinai oscillator trap

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We study numerically the evolution of Bose-Einstein condensate in the Sinai oscillator trap described by the Gross-Pitaevskii equation in two dimensions. In absence of interactions this trap mimics the properties of Sinai billiards where the classical dynamics is chaotic and the quantum evolution is described by generic properties of quantum chaos and random matrix theory. We show that, above a certain border, the nonlinear interactions between atoms lead to emergence of dynamical thermalization which generates the statistical Bose-Einstein distribution over eigenmodes of system without interactions. Below the thermalization border the evolution remains quasi-integrable. Such a Sinai oscillator trap, formed by the oscillator potential and a repulsive disk located in center vicinity, had been already realized in first experiments with the Bose-Einstein condensate formation by Ketterle group in 1995 and we argue that it can form a convenient test bed for experimental investigations of dynamical origins of thermalization. Possible links and implications for Kolmogorov turbulence in absence of noise are also discussed.

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I. INTRODUCTION

One of the first experimental realizations of Bose-Einstein condensate (BEC) has been done with sodium atoms trapped in a novel trap that employed both magnetic and optical forces [1]. In this trap, the repulsive optical potential is created by tightly focusing an intense blue-detuned laser that generates a repulsive optical plug bunging a hole in a center of magnetic trap where nonadiabatic spin flips lead to a loss of atoms. Further developments of BEC traps, a remarkable progress of BEC experiments and theory of BEC in trapped gases are reviewed in [2–5].

In spite of these achievements, the fundamental question about interplay of dynamics, interactions and thermalization of BEC in a concrete trap configuration still waits its clarification. In this work we address this question in the frame of the Gross-Pitaevskii equation (GPE) [4, 5] for the two-dimensional (2D) version of the trap used in the experimental setup [1]. Thus the trap potential is represented by a 2D harmonic potential and a rigid disk which center is located in a vicinity of the center of harmonic potential. If the harmonic potential is replaced by rigid walls forming a square or rectangle then the classical dynamics in such a Sinai billiard is proven to be completely chaotic [6]. The recent analysis of the trap with the walls formed by a harmonic potential shows that the dynamics remains chaotic with a very small measure of integrable dynamics [7]. Due to a similarity with a Sinai billiard such a trap was called the Sinai oscillator [7]. Since the realization of rigid walls is rather difficult for experimental realization the case of Sinai oscillator trap becomes much more attractive for combined theoretical and experimental investigations. In fact a Sinai oscillator trap in three-dimensions (3D) had been implemented in [1]. Here we restrict our investigations to the 2D case expecting that its main features will be preserved in 3D.

The quantum properties of Sinai oscillator are characterized by the Wigner-Dyson statistics of energy levels [8] typical for the random matrix theory [9]. The properties of the eigenstates are typical for those of systems of quantum chaos and now are well understood (see e.g. [10, 11]).

Below we show that the Sinai oscillator trap in 2D is also characterized by the properties of quantum chaos: the quantum eigenstates of Sinai oscillator are ergodic and the level spacing statistic is described by the random matrix theory, in agreement with the Bohigas-Giannoni-Schmit conjecture [8, 10]. However, still there is no thermalization in this system since the eigenstates are preserved in absence of interactions. Thus, in this work we analyze the dynamics and thermalization conditions for BEC in the Sinai oscillator in the frame of the GPE equation. The GPE description is valid in the regime when the BEC temperature T is below the critical temperature of Bose-Einstein condensation T_c [12] and when the validity of GPE description is well justified [4, 5].

Even from the mathematical view point the question about existence of the solution of GPE in such a trap, with chaotic classical ray dynamics, at moderate nonlinearity and large times remains an open problem (see e.g. [13, 14]). Indeed, the GPE can be rewritten in the basis of linear eigenstates (modes) where the coupling between modes takes place only due to the nonlinearity in GPE. In this representation each mode can be considered as an independent oscillator degree of freedom and in case of thermalization, induced by nonlinearity, one should expect energy equipartition over all modes [12] that leads to ultra-violet catastrophe and energy transfer to high energy modes. In fact, the Planck constant and the Planck law had been introduced for a black-body radiation to avoid such a divergence [15]. However, the Planck distribution is valid for quantum systems while in our case of the GPE Sinai oscillator there is no second quantization. Thus due to classical nonlinear interactions between modes one would expect to have a classical ergodicity with equipartition of energy between modes.

Indeed, such an equipartition expectation was at the origin of the studies of the Fermi-Pasta-Ulam (FPU) problem [16, 17]. Its absence was attributed to location of initial conditions below chaos border and system proximity to the integrable Toda lattice (see e.g. [18, 19] and Refs. therein). Thus the FPU oscillator chain has certain specific features which break system ergodicity in energy. However, it is natural to expect that in a generic case, when eigenstates of a linear system are ergodic and dynamical chaos of classical trajectories takes place, the energy equipartition over modes should appear above certain border of nonlinear interaction strength between modes.

In spite of these expectations of energy equipartition over modes, the recent studies of the GPE in Bunimovich stadium showed that the nonlinearity produces an effective dynamical thermalization in a completely isolated system, without any contact with external thermal bath, with the probabilities over linear modes described by the Bose-Einstein (BE) distribution [7]. Thus the probabilities on high energy modes drop rapidly and the ultraviolet catastrophe is absent. An experimental realization of the Bunimovich billiard with cold atoms is possible but is not so simple. Due to this reason we consider here the GPE Sinai oscillator trap which in fact has been already built in [1-3] but without investigation of phenomenon of dynamical thermalization. Our results show the presence of dynamical thermalization with BE distribution in this system even if some aspects still should be clarified for time evolution of very large time scales.

The model description, the quantum chaos features of linear system are described in Section II, the thermalization equations and formalism are presented in Section III; the obtained numerical results for the GPE evolution are presented and discussed in the Section IV; numerical methods and behavior on large time scales are discussed in Section V; the discussion of main results is presented in Section VI.

II. MODEL DESCRIPTION AND QUANTUM CHAOS PROPERTIES

The dynamics of the classical Sinai oscillator is described by the Hamiltonian:

$$H = \frac{1}{2m}(p_x^2 + p_y^2) + \frac{m}{2}(\omega_x^2 x^2 + \omega_y^2 y^2) + V_d(x, y) \quad , \quad (1)$$

where the first two terms describe 2D oscillator with frequencies ω_x, ω_y and the last term describes the potential



wall of elastic disk of radius r_d centered at (x_d, y_d) . In our studies we fixed the mass m = 1, frequencies $\omega_x = 1$, $\omega_y = \sqrt{2}$ and disk radius $r_d = 1$. The disk center is placed at $(x_d, y_d) = (-1/2, -1/2)$ so that the disk bungs a hole in the center as it was the case in the experiments [1].

It is convenient to characterize the classical dynamics by the Poincaré section using the canonical variables taken at the moment of time when a particle bounces with the disk. At that moment we take the phase φ , given by the angle measured from x-axis, and the conjugated dimensionless orbital momentum $\kappa = \sin \theta$, where θ is the angle of momentum \vec{p} counted from the normal to the circle (see Fig. 1). Such a pair of conjugated variables represents a standard choice for description of dynamics in billiards (see e.g. [10, 11]).

The results for Poincaré sections at Fig. 1 show that almost all phase space, accessible for dynamics at a given energy, is chaotic (see e.g. [20, 21] on properties of dy-





FIG. 2: (Color online) Nearest-neighbor spacing distribution P(s) for the first 2500 unfolded eigenenergies of the Sinai oscillator (1). The red dashed curve represents the Wigner surmise $P(s) = (\pi s/2) \exp(-\pi s^2/4)$. Insert panel shows energy eigenvalues $E = E_m$ as a function of m for the first 50 eigenvalues $1 \le m \le 50$. Dashed blue curve represents the theoretical Weyl law $m(E) = E^2/(2\sqrt{2}) - E/2$.

namical chaos). Only very tiny stability islets of regular motion can be found at E = 3 (practically not visible on the Poincaré section). At small energy E = 1.5 the dynamics is located only on one side of the disk (variation of φ is bounded, $-0.4 < \varphi/\pi \le 1$) due to symmetry breaking of the shift of disk center. At larger energies the trajectories make complete rotations around the disk. The amplitude of oscillations grows with energy approximately in the same way as in a usual 2D oscillator in absence of disk. However, the scattering on disk makes the dynamics chaotic in a similar manner as for a standard Sinai billiard [6].

The BEC evolution in the Sinai oscillator trap is described by the GPE, which reads:

$$i\hbar \frac{\partial \psi(\vec{r},t)}{\partial t} = -\frac{\hbar^2}{2m} \vec{\nabla}^2 \psi(\vec{r},t)$$

$$+ \left[\frac{m}{2} (\omega_x^2 x^2 + \omega_y^2 y^2) + V_d(x,y) \right] \psi(\vec{r},t)$$

$$+ \beta |\psi(\vec{r},t)|^2 \psi(\vec{r},t) .$$
(2)

Here in (2), we use the same oscillator and disk parameters as in (1) and take $\hbar = 1$. The wave function is normalized to unity $W = \int |\psi(x, y)|^2 dx dy = 1$. Then the parameter β describes the dimensionless nonlinear interactions of atoms in BEC.

Since the classical dynamics is chaotic and the measure of integrable islands is very small it is natural to expect that at zero nonlinearity $\beta = 0$ the Sinai oscillator belongs to systems of quantum chaos [8, 10, 11]. Indeed, using the advanced methods of quantum chaos on numerical computation of eigenenergies and eigenstates in chaotic billiards (see e.g. [22]), we find numerically several thousands of eigenenergies E_m and eigenstates m(linear modes) at $\beta = 0$. The level spacing statistics P(s) for the first 2500 energy levels with the unfolding procedure (see e.g. [10]) is shown in Fig. 2. The results are in good agreement with the Wigner surmise confirming the validity of the Bohigas-Giannoni-Schmit conjecture for the Sinai oscillator trap [8, 10]. The system energy E_m grows with the level number m in agreement with the Weyl law as it is shown in the insert of Fig. 2.

The linear eigenstates ϕ_m have a rather complex structure covering the accessible area in (x, y) plane with chaotic fluctuations. The probability distributions in (x, y) plane are shown for first 100 eigenstates in [23]; some eigenstates are also shown in Figs. below.

The GPE (2) can be also rewritten in the basis of linear eigenstates ϕ_m using the completeness of this basis and presenting the wave function by expansion $\psi(x, y, t) = \sum_m C_m(t)\phi_m(x, y)$, where $C_m(t)$ are time dependent probability amplitudes in this basis. Then in this basis the GPE reads:

$$i\frac{\partial C_m}{\partial t} = E_m C_m + \beta \sum_{m_1 m_2 m_3} U_{mm_1 m_2 m_3} C_{m_1} C_{m_2}^* C_{m_3} .$$
(3)

Here the transitions between eigenmodes appear only due to the nonlinear β -term and the transition matrix elements are

$$U_{mm_1m_2m_3} = \int dx dy \,\phi_m^* \phi_{m_1} \phi_{m_2}^* \phi_{m_3} \,. \tag{4}$$

In our case, in absence of a magnetic field, the eigenstates ϕ_m are real, but we keep the general expression valid also for complex eigenstates.

A similar type of representation (4) was used for the analysis of effects of nonlinearity on the Anderson localization in disordered lattices [24] known also as the DANSE model [25, 26]. In this model it was found that a moderate nonlinearity leads to a destruction of the Anderson localization of linear eigenmodes and a subdiffusive spreading of wave packet over lattice sites with time. Such a spreading has been studied by different groups (see e.g. [24–28] and Refs. therein). Even if the representations for the GPE Sinai oscillator (4) and DANSE models are similar there are significant differences: (a) in DANSE the eigenenergies are bounded in a finite energy band while here $E_m \propto \sqrt{m}$ are growing with m (we note that for the Bunimovich billiard we have $E_m \propto m$ [7]); (b) in DANSE the transitions $U_{mm_1m_2m_3}$ give coupling mainly between states inside the same localization length while here there are transitions even between vary different m values. At the same time we should say that the properties of matrix elements $U_{mm_1m_2m_3}$ are still waiting their detailed analysis for the cases of Sinai oscillator and Bunimovich billiard.

We also note that the question of energy transfer to high energy modes has certain links with the Kolmogorov turbulence which is based on the concept of energy flow from large to small scales via the inertial interval (see [29, 30] and Refs. therein). In this concept energy is injected at large scale and absorbed on small scales and a presence of some small noise is assumed to induce thermalization flow. In such an approach the "quantum" turbulence in GPE (or nonlinear Schödinger equation (NLS)) has been studied in a rectangular 2D billiard [31] and in 3D cube [32]. Due to the billiard shape chosen there, the ray dynamics is integrable and it is not obvious if the dynamical thermalization takes place in such a billiard in absence of noise. Below we will see that at moderate nonlinearity and absence of noise there is no dynamical thermalization in oscillator trap and billiards of rectangular shape (see the later case in [7]). In fact, in purely dynamical systems (without external noise) it is possible that the Kolmogorov flow to high modes can be stopped by KAM integrability and Anderson localization [33].

Our results show that the quantum chaos for linear eigenmodes facilitates onset of dynamic thermalization, appearing in an isolated system without any external noise, when the strength of nonlinear term is above a certain dynamical thermalization border $\beta > \beta_c$.

III. THERMODYNAMIC FORMALISM

The dynamical thermalization in nonlinear chains with disorder has been studied in [34, 35] where it was shown that the quantum Gibbs distribution appears in an isolated system above a certain border of nonlinearity $\beta > \beta_c$. The dynamical thermalization for the GPE in a chaotic Bunimovich billiard has been established in [7].

Below for a reader convenience we present the thermalization formalism which directly follows from the standard results of statistical description of Bose gas [12]. Indeed, we assume that the nonlinearity is moderate and that the nonlinear term gives only a small energy shift which can be neglected. Then the energy levels are those of the quantum Sinai oscillator with the usual quantum chaos properties and the energy levels E_m at $\beta = 0$. Since the energy and the norm of the system (2) are conserved the thermalization ansatz gives the steady-state probabilities ρ_m on energy levels:

$$\rho_m = 1/[\exp[(E_m - E_g - \mu)/T] - 1], \qquad (5)$$

where where $E_g = 1.685$ is the energy of the ground state, T is the temperature of the system, $\mu(T)$ is the chemical potential dependent on temperature. The parameters T and μ are determined by the norm conservation $\sum_{m=1}^{\infty} \rho_m = 1$ (we have only one particle in the system) and the initial energy $\sum_m E_m \rho_m = E$. The entropy S of the system is determined by the usual relation $[12]: S = -\sum_m \rho_m \ln \rho_m$. The relation (5), with normalization condition and equation of energy, determines the implicit dependencies on temperature $E(T), S(T), \mu(T)$.

As it is pointed in [7, 34, 35], the advantage of energy E and entropy S is that both are extensive variables, thus they are self-averaging and due to that they have reduced fluctuations. Due to this feature S and E are especially convenient for verification of the thermalization ansatz



FIG. 3: (Color online) Dependence of entropy S on energy E, obtained from the GPE time evolution (2) for initial states taken as first 50 eigenstates ϕ_m of the quantum Sinai oscillator. Blue (black) and red (gray) symbols show the cases of nonlinearity $\beta = 0.5$ and $\beta = 4$ respectively, while circles and crosses represent the system with and without the elastic disk r_d respectively. The entropy S is computed from $\rho_m = \langle |\langle m | \psi \rangle|^2 \rangle_t$ averaged over time intervals $t \in [500, 1500]$ (top panel), and $t \in [1500, 2500]$ (bottom panel). The dashed curve shows the theoretical thermalization ansatz of Bose-Einstein distribution (5).

(5) which gives the theoretical dependence S(E) in the assumption that the dynamical thermalization emerges in the GPE Sinai oscillator due to dynamical chaos in absence of any external noise and thermostat coupling.

IV. NUMERICAL RESULTS

The numerical integration of GPE (2) is done following the approach used for the Bunimovich billiard in [7]: we introduce the space grid with a usual size of $N_s = n_x \times n_y = 201 \times 141 = 28341$ spacial points. The time step is performed with the Trotter decomposition of linear and nonlinear terms with a usual time step $\Delta t = 0.01$. Thus, the nonlinear term gives the wave function transformation in coordinate space $\overline{\psi}(x,y) = \exp(-i\Delta t\beta |\psi(x,y)|^2)\psi(x,y)$, which is then transformed from coordinate space to the linear eigenbasis ϕ_m . The transformation from space grid to linear eigenfunction index m is done via a precomputed transfer matrix A_{jm} (here $1 \leq j \leq N_e$ is an index of space grid). After that the linear propagation step is performed with expansion coefficients in the eigenstate basis $C(t + \Delta t) = \exp(-i\Delta t E_m)C_m(t)$. Then the back transformation from linear basis ϕ_m to coordinate basis



FIG. 4: (Color online) Dependence of temperature T and chemical potential μ on energy shown in top and bottom panels respectively. Black curves represent the theoretical ansatz given by the Bose-Einstein distribution (5), while red circles represent numerical data $T = (T_1(E) + T_2(S))/2$ and $\mu = (\mu_1(E) + \mu_2(S))/2$ ($T_{1,2}$ and $\mu_{1,2}$ values are computed from E and S respectively) for initial states given by first 50 linear eigenstates and probabilities $\rho_m = \langle |\langle m | \psi \rangle|^2 \rangle_t$ averaged over interval $t \in [500, 1500]$. Here $\beta = 4$.

finally gives $\psi(t + \Delta t)$. We use usually $N_e = 2000$ linear eigenstates and the grid with $N_s = 28341$ points for such a time evolution. In addition to the space representation $\psi(x, y, t)$ we also compute the wave function in the momentum representation using the standard relation $\phi(\vec{p}, t) = \int \psi(\vec{r}, t) \exp(-i\vec{r}\vec{p}/\hbar) dr^2/2\pi\hbar$ with a discrete Fourier transform. The time evolution computed in this way gives an approximate energy E and norm W conservation. Typically we have at time t = 2000 the variation of these integrals being $\delta W/W = 0.001(0.004)$ and $\delta E/E = 0.002(0.004)$ for initial state at m = 10(40) respectively. We return to a more detailed discussion of the accuracy of computations in Section V.

During the time evolution we determine the probabilities of wave function in the linear eigenmodes $\rho_m = \langle |\langle m | \psi \rangle|^2 \rangle_t = \langle |C_m|^2 | \rangle_t$ averaged over a certain time interval. Usually we choose this time interval as approximately last half (or similar to that) of the whole evolution range to obtain approximate steady-state values of ρ_m . From these averaged values we determine the entropy $S = -\sum_m \rho_m \ln \rho_m$. Thus starting from different initial states, chosen as linear eigenstates ψ_m , we obtain numerically the dependence S(E) which is compared with the prediction of the Bose-Einstein thermalization ansatz



FIG. 5: (Color online) Spacial probability distributions $|\psi(x, y)|^2$ for the GPE Sinai oscillator. Panels a) and b) show the linear eigenstates m = 1 (ground state) and m = 24 with eigenergies $E_1 = 2.417$ and $E_{24} = 9.16$ respectively. Panels c) and e) have the initial state m = 24 of panel b) and show the average distributions at long times with averaging over large interval $t \in [1500, 2500]$ for $\beta = 0.5$ and $\beta = 4$ respectively. Panel d) shows the average distribution for short time interval (snapshot) $t \in [2000, 2005]$ for $\beta = 4$. Panel f) shows the thermal Bose-Einstein distribution (5) for energy $E_{24} = 9.16$ (to be compared with panel e)). Probability is shown by color bar changing from zero (black) to maximum (yellow/gray).

The comparison of numerical data with the theoretical curve obtained from (5) is shown in Fig. 3. It is clear that at small $\beta = 0.5$ the nonlinear term leads to excitation of certain eignemodes of the Sinai oscillator but the numerical data for S(E) are pretty far from the theoretical dashed curve given by the Bose-Einstein distribution (5). For the 2D oscillator without disk the excitation is significantly weaker then for the Sinai oscillator with S(E) values being very far form the theory both for $\beta = 0.5$; 4. In contrast to that, for the GPE Sinai oscillator at $\beta = 4$ our numerical data for S(E) are close to the theoretical thermalization ansatz. For the time interval $t \in [500, 1500]$ (Fig. 3 top panel) the obtained S values for $35 < m \le 50$ (10.5 < E < 13 are somewhat below the theoretical curve. We attribute this to the fact that for large m the effective amplitude of nonlinear term in (2) is reduced $|\psi|^2 \sim 1/x^2 \sim 1/E \sim 1/\sqrt{m}$ and hence, it takes a longer time for dynamical chaos to establish dynamical thermalization. Indeed, the nonlinear energy shift $\delta E_{\beta} \sim \beta |\psi|^2 \sim \beta/\sqrt{m}$ so that the thermalization time t_T should be at least proportional to $t_T \sim 1/\delta E_{\beta} \sim \sqrt{m}/\beta$. Indeed, at large times $t \in [1500, 2500]$ (Fig.3 bottom panel) we find the values of S being significantly more close to the thermalization ansatz for $35 < m \leq 50$ (10.5 < E < 13). Thus the results of Fig. 3 show the onset of dynamical thermalization for moderate values of $\beta > \beta_c \sim 1$.



FIG. 6: (Color online) Momentum probability distributions $|\psi(p_x, p_y)|^2$ shown for the same cases as in 6 panels of Fig.5. Probability is shown by color bar changing from zero (black) to maximum (yellow/gray).

As for the case of Bunimovich stadium [7], we expect that the thermalization border $\beta_c \sim 1$ (definitely $0.5 < \beta_c < 4$) is independent of m. Indeed, the nonlinear energy shift $\delta E_{\beta} \sim \beta/\sqrt{m}$ and the level spacing $\Delta E \sim 1/\sqrt{m}$ scale with m in a similar way (see Fig. 2 for the dependence E_m) so that we can expect chaos and thermalization to be set in at $\delta E_{\beta} > \Delta E$, thus leading to $\beta_c \sim 1$. Indeed, similar estimates have been confirmed in systems of coupled nonlinear oscillators [24, 35, 36]).

Another confirmation of the onset of dynamical thermalization is shown in Fig. 4. Indeed, according to (5) the temperature can be determined from an initial energy E, giving $T_1(E)$, or from an average value of S, giving $T_2(S)$. In a similar way we can determine $\mu_1(E)$ and $\mu_2(S)$. The dependence of average numerical values $T = (T_1 + T_2)/2$, $\mu = (\mu_1 + \mu_2)/2$ on energy E are shown in Fig. 4 being in a good agreement with the thermalization ansatz (5). Again, we attribute certain deviations for $35 < m \leq 50$ to not sufficiently large times reached in numerical simulations.

The transition from nontermalized (quasi-integrable) regime to dynamical thermalization is also visible from the spacial probability distributions shown in Fig. 5. We start from a typical initial state m = 24 shown in panel (b). For $\beta = 0.5 < \beta_c$ a snapshot distribution at t = 2000 (panel c)) remains very similar to the initial state showing the absence of thermalization and dominance of initial mode. In contrast to that for $\beta = 4 > \beta_c$ a snapshot at t = 2000 (panel d)) shows that the distribution have a dominant component at the ground state mode shown in panel a). The distribution averaged over a large time interval, assumed to be close to a steady-state, is shown in panel f). It is indeed very similar to the theoretical steady state probability distribution $|\psi_{st}(x,y)|^2 = \sum_m \rho_m |\phi_m(x,y)|^2$ where ρ_m are given by the Bose-Einstein distribution (5) (see Fig. 5 panel e)).

The probability distributions in the momentum space (p_x, p_y) , corresponding to cases of Fig. 5, are shown in Fig. 6. These data also show a clear absence of thermalization for $\beta = 0.5$ (panels b), c)) and close similarity between the theoretical distribution (panel e)) and average distribution (panel f)).



FIG. 7: (Color online) Time evolution of probabilities $\rho_m(t)$ in the basis of linear eigenmodes for the initial state m = 24 at $\beta = 4$. The probabilities $\rho_m(t)$ are averaged over time $\delta t = 10$ to reduce fluctuations. Color bar shows probabilities from zero (black) to maximum (white).

Thus the results of this Section provide a good confirmation of onset of dynamical thermalization at moderate nonlinearity $\beta \approx 4 > \beta_c \sim 1$. However, it is also important to analyze the larger scale evolution on times being larger than those considered here with $t \leq 2500$. This consideration is presented in next Section.

V. LARGE TIME SCALES AND NUMERICAL METHODS

The question about long time scale evolution of system (2) requires further extensive studies with improved accuracy on numerical simulations. Indeed, at times t > 2500 we see a tendency of probability accumulation at the ground state ϕ_1 of linear system. First signs of this trend are seen in Fig. 7 showing evolution $\rho_m(t)$ with appearance of large values of ρ_1 at $t \approx 2300$.



FIG. 8: (Color online) Time evolution of probabilities $\rho_m(t)$ in the basis f of linear eigenmodes for the initial state m = 24at $\beta = 4$. The probabilities $\rho_m(t)$ are averaged over time $\delta t = 10$ to reduce fluctuations. Ten panels show $\rho_m(t)$ for m = 1, 2, 3, 4, 5, 6, 7, 8, 9 and m = 24; red dashed lines show the theoretical values of ρ_m from the thermalization ansatz of Bose-Einstein distribution (5).

A more detailed view of dependence of $\rho_m(t)$ on time t for several selected m values is presented in Fig. 8 up to t = 5000. For m > 1 there are large fluctuations in time and it is clear that averaging on large time scales is required to obtain statistically stable values of ρ_m . For the extensive variables like energy E and entropy S these fluctuations are reduced and that is the reason due to which the data for the curve S(E) are less fluctuating. However, for m = 1 in Fig. 8 there is a steady growth with a apparent saturation at $t \approx 4500$ at the value $\rho_1 \approx 0.4$ which is by a factor 10 larger than the theoretical value shown by the dashed line. The spacial distribution of probability at such large times also demonstrates a strong accumulation of probability at the ground state as it is shown in Fig. 9. A video of evolution on large times is available at [23]. A similar probability accumulation at



FIG. 9: (Color online) Long time average probability distribution $\langle |\psi(x,y)|^2 \rangle_t$ shown in coordinate space. The evolution starts from the initial linear eigenstate m = 24 (see Fig. 5), the average is done over the time interval $t \in [4000, 5000]$; here $\beta = 4$ (to compare with panels of Fig. 5 with same colors).

the ground state is seen for other initial states with $\psi(t = 0) = \phi_{40}; \phi_{60}$. We note that we had no such accumulation of probability in the ground state in the studies of GPE Bunimovich billiard [7].

We also considered the time evolution for a different initial state taken as initial coherent wave packet centered at certain position (x_0, y_0) (of course the packet is only approximately Gaussian since the probability is zero at the disk border). The initial distribution and snapshots at a few moments of time are shown in Fig. 10 (panels a), b), c)). The time evolution video is available at [23] for $\beta = 0$ and $\beta = 4$ on short and long time scales. The probability averaged over time interval $t \in [4000, 5000]$ is shown in panels f) and d) $\beta = 0$ and $\beta = 4$ respectively. The case with interactions shows a tendency of accumulation of probability at low m modes in a qualitative agreement with the thermal ansatz distribution shown in panel e). However, it is visible that the GPE case has a larger probability on low energy modes m. In contrast, the case with $\beta = 0$ (panel f)) has large probability in initial high modes m.

The average probability distributions over linear modes are shown at large times in Fig. 11. For the initial state in a form of coherent packet there is a clear displacement of highest probabilities from initial energies $E_m \approx 14.7$ to modes with m = 1, 2, 4. In this case there is no strong accumulation of probability at the ground state. In fact, the average value of $\rho_1 \approx 0.22$ is comparable with the thermalization ansatz value. Also for this state with energy $E \approx 14.7$ we find numerically the average entropy value S = 4.96 being close to the theoretical value of (5) with S = 5.47. However, the fluctuations of average ρ_m probabilities are too strong and the comparison with the thermalization ansatz curve is only qualitative. For the initial state with ϕ_{24} the probability



FIG. 10: (Color online) Space probability distributions $|\psi(x, y)|^2$ for the case of an initial state given by a coherent wave packet centered at $(x_0, y_0) = (3, 3)$ (and zero probabilities on the disk). Panel a) shows the initial state distribution with E = 14.707; panels b) and c) show distributions with $\beta = 4$ at t = 2.9 and t = 8 respectively. Long time averages over the interval $t \in [4000, 5000]$ are shown in panel d) and f) for $\beta = 4$ and $\beta = 0$ (linear evolution) respectively. Panel e) shows the theoretical space probability from the Bose-Einstein distribution (5) for energy E = 14.707.

 ρ_1 is significantly larger than the theoretical value (by a factor 10), then the decay of ρ_m with E_m approximately follows the thermalization ansatz but also the fluctuations are larger. We think that the fluctuations are larger for the case of coherent state due to a larger number of initially exited linear eigenmodes compared to the case with $\psi(t=0) = \phi_{24}$. Also the initial energy of the coherent state is E = 14.71 being larger than energy of m = 24 with E = 9.16 and hence longer times are required for complete thermalization of the coherent state.

To check the validity of numerical integration for large time scales we performed a number of checks shown in Fig. 12 and Fig. 13. With this aim we varied the integration time step Δt , the number of lattice points N_s in the coordinate space $(-x_{min}, x_{max}; -y_{min}, y_{max})$ and the



FIG. 11: (Color online) Probability ρ_m vs. energy of linear eigenstate E_m . Top panel shows the case of initial state given by a coherent state localized at $(x_0, y_0) = (3, 3)$ from Fig.10; bottom panel shows the case of initial linear eigenstate m =24 (semi-logarithmic scale). Black circles and blue squares show ρ_m for initial state and for long time average with $\beta = 4$ and $t \in [4000, 5000]$ respectively. Solid red line shows the total energy, while dashed orange line represents the theoretical distribution ρ_m (5) for corresponding energy E = 14.71 and E = 9.16 (top and bottom panels respectively).

spacial range of the lattice determined by these min/max values of x, y. The integration scheme gives a slow decrease of norm W and energy E with time indicating that there are some effective dissipation induced by numerical integration. These checks show that the most sensitive parameter is the number of eigenstates N_e used in the transformation matrix $A_{j,m}$ from coordinate space to linear eigenmodes. Up to time $t \approx 800$ the all numerical curves in Fig. 13 give the same results showing the validity of numerical integration. However, for $N_e = 1000$ the non-conservation of W and E becomes significant at large times giving significantly different values of ρ_1, ρ_4 at $t \approx 1000 - 1500$ for $N_e = 1000$ and $N_e = 2000$. Thus we use the numerical parameters with $N_e = 2000$ for all results presented in above Figs. 1- 11 since this choice provides the best conservation of W and E.

The above verification show that our integration scheme introduces a hidden intrinsic effective dissipation and due to this reason we consider that this can be the reason of probability accumulation at the ground state (or a few states with low m values) at large times. Due to this reasons we consider that our numerical results are reliable only up to finite times $t \leq t_{num} \approx 2500$. We believe that better symplectic integration schemes should be invented to allow to extend numerical studies of system (2) at larger time scales.



3000

4000

5000

FIG. 12: (Color online) Time evolution of energy and norm conservation for different parameters of numerical simulations with the initial linear eigenstate m = 24 and $\beta = 4$. Top panel and bottom panel show the evolution of $\Delta E = E - E_{24}$ and $\Delta W = 1 - \sum_{m} \rho_m$. The simulation parameters are: $N_e = 2000, \Delta t = 0.01$ and number of spacial lattice of $N_s = 201 \times 141$ points in the rectangle given by $(x_{\min}, y_{\min}) = (-\sqrt{234} \approx -15.3, -\sqrt{117} \approx -10.82)$ and $(x_{\max}, y_{\max}) = (\sqrt{234} \approx 15.3, \sqrt{117} \approx 10.82)$ (black solid curves); $N_e = 1000, \Delta t = 0.01$ with $N_s = 201 \times 141$ (red dashed curves); $N_e = 1000, \Delta t = 0.005$ with $N_s = 201 \times 141$ (green dot-dashed curves); $N_e = 1000, \Delta t = 0.005$ with $N_s = 201 \times 141$ (green dot-dashed curves); $N_e = 1000, \Delta t = 0.005$ with $N_s = 361 \times 255$ lattice points in the same region (blue dotted curves); red and blue curves practically coincide in the bottom panel.

2000

1000

VI. DISCUSSION

The results of this work demonstrate the emergence of dynamical thermalization of BEC described by the GPE equation of wave function time evolution in Sinai oscillator trap. The classical dynamics of rays in such a trap is chaotic and the quantum properties of this system in absence of interactions are described by well known results of the field of quantum chaos. The dynamical thermalization appears above a certain critical strength of interaction $\beta > \beta_c$. In this thermalized phase the probability distribution ρ_m over linear eigenmodes (at $\beta = 0$) is well described by the standard statistical Bose-Einstein distribution. We stress that the dynamical thermalization appears in a completely isolated system without any external noise. We point that this thermal distribution is drastically different from energy equipartition over modes which is usually expected for nonlinear oscillator lattices, including the Fermi-Pasta-Ulam problem, leading to the ultra-violet catastrophe. Thus our results show that energy is redistributed only over certain low energy modes and that there is no energy flow to high energy modes.



FIG. 13: (Color online) Time evolution of probabilities $\rho_m(t)$ for parameters of Fig. 12 shown for m = 1, 4, 6, 24 with the same colors as in Fig. 12.

We think that this result may have interesting implications to the dynamical consideration of Kolmogorov turbulence which assumes the presence of energy flow from large (low energy) to small (high energy) spacial scales in presence of noise [29, 30]. Our results indicate that in absence of noise such energy flow can be absent due to absence of energy equipartition. Of course, further investigations of this system in numerical simulations on larger time scales are highly desirable in view of numerical difficulties discussed in the previous Section.

The trap configuration considered here had been already realized experimentally in 3D [1] and we believe that further experimental investigations of dynamical thermalization in the Sinai oscillator trap are within reach of modern experiments with cold atoms and BEC. The variation of interaction strength between atoms by means of Feshbach resonance should be able to detect a transition from quasi-integrable phase at $\beta < \beta_c$ to the phase of dynamical thermalization at $\beta > \beta_c$. The case of Litium 6 atoms, where the interactions can be changed in a broad range (see e.g. [37]), can be a good test bed for the studies of dynamical thermalization in the Sinai oscillator trap and fundamental origins of thermalization in isolated systems.

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